PHYS 575A/B/C Autumn 2015

Radiation and Radiation Detectors

Course home page:

http://depts.washington.edu/physcert/radcert/575website/

9: Case studies: Non-Cherenkov neutrino detectors; neutron detectors; accelerators

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Course calendar (revised)

week	date	day	topic	text		
1	10/1/15	Thurs	Introduction, review of basics, radioactivity, units for radiation and dosimetry	Ch. 1, notes		
2	10/6/15	Tues	Radioactive sources; decay processes; Ch. 1, notes			
3	10/13/15	Tues	Photomultiplier tubes and scintillation counters; Counting statistics	Chs. 3, 8, 9 (I-V)		
3	10/15/15	Thurs	LAB: Room B248 Scopes, fast pulses; PMTs and scintillation counters; standard electronics modules	Chs. 4, 9, 16, 17		
4	10/20/15	Tues	Overview of charged particle detectors	Ch. 4		
4	10/22/15	Thurs	LAB: Room B248 Coincidence techniques; nanosec time measurement, energy from pulse area	Chs. 17, 18		
5	10/27/15	Tues	Interaction of charged particles and photons with matter; counting statistics; gas detectors; Proposal for term paper must be emailed to JW by today	Chs. 2, 3; Chs. 5, 6, 7		
6	11/3/15	Tues	ionization chambers; solid-state detectors Chs. 11, 12, 13			
7	11/10/15	Tues	Statistics for data analysis; Case studies: classic visual detectors (cloud and bubble chambers, nuclear emulsion, spark chambers)			
8	11/17/15	Tues	Case studies: Cosmic ray detectors (Auger, Fermi gamma ray observatory); Cherenkov detectors: atmospheric Cherenkov, triggering Cherenkov			
9	11/24/15	Tues	Case studies: neutrino detectors (IceCube, Daya Bay, Majorana), Detecting neutrons; high energy accelerators; Ch. 19, notes Ch. 14, 15, 18			
10	12/1/15	Tues	Finish case studies; begin student presentations Notes			
11	12/8/15	Tues	Student presentations -			
11	12/10/15	Thurs	Student presentations			

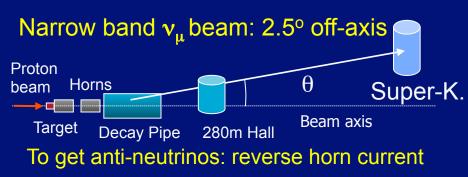
Tonight

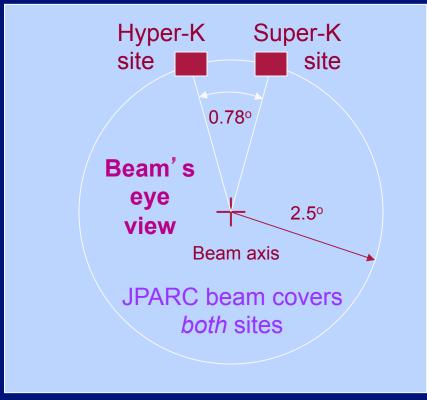
Announcements

- Presentation dates: Tues Dec 1, Tues Dec 8, and Thurs Dec 10
 - You MUST send me your presentation (pdf or ppt) no later than 5:30 pm on the day of your talk
 - I will upload all slides for each session so online attendance is possible

Please send me you	r presentation pp	t/pdf (or URL) at least 1 hou	r before class on your date
Day	Time	Name	Topic
12/1/2015	7:00 PM	Per Provencher	Low Background Laboratories
	7:20 PM	Rick McGann	Neutron Generation and Effects on Materials and Electronic
	7:40 PM	Chris Provencher	Bremsstrahlung
	8:00 PM	Charles Ko	Radiometric Dating
	8:20 PM		
12/8/2015	6:40 PM	Diana Thompson	NORM
	7:00 PM	Shawn Apodaca	Fast Neutron Time of Flight and Spectroscopy
	7:20 PM	Erin Board	Cosmic Radiation and Shielding
	7:40 PM	Louie Cueva	Thermal Neutron Detection
	8:00 PM	Xavier Garcia	Silicon PMTs
	8:20 PM	Padmaja Vrudhula	Dosimetry
12/10/2015	6:40 PM	Nathan Hicks	Methods of Radionuclide Production for Medical Isotope Usability: Meeting the Demand
	7:00 PM	Farrah Tan	QCD
	7:20 PM	Nicolas Michel-Hart	microXRF
	7:40 PM	Michael Esuabana	proton-Boron11 fusion
	8:00 PM	Kaifu Lam	Synchrotrons
	8:20 PM	Johnathan Slack	X-rays/Gamma rays of comets and asteroids

JPARC neutrino beam uses "off-axis" technique





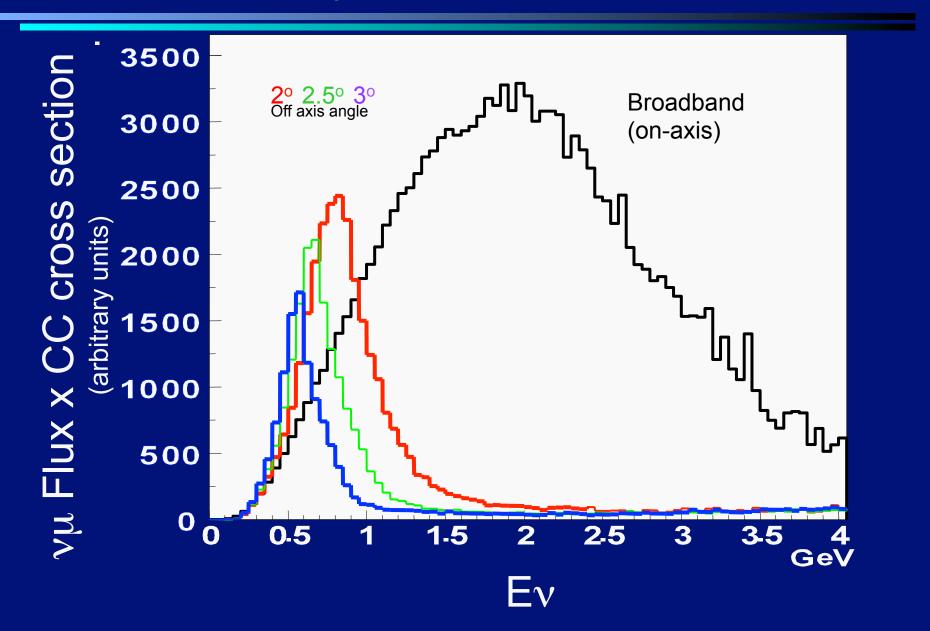
T2K-I event statistics at SK: (Off axis-2.5 deg, 22.5 kt, per year)

- 2200 total ν_μ events
- 1600 v_u CC
- $v_e \sim 0.4\%$ at v_u peak E



Target/Horn magnet – test setup Must handle pulses of 100s of kA

T2K Neutrino E spectra: broadband vs off-axis



Low-energy (reactor) neutrino oscillation experiments

- Nuclear reactors provide high fluxes of electron antineutrinos
 - Flux directly related to power output of nuclear plant
 - Baseline L can be shorter since energy is few MeV (recall: L/E ~ proper time for neutrino)
- Pure water detectors are not optimal small cross-section*
- Use liquid scintillator (fluors in mineral oil) or Cd compounds
 - * But: Doping water with gadolinium produces huge cross sections
 - GADZOOKS = plan to turn Super-K into a Gd detector for antinu
 - Neutrons captured by Gd, 90% capture efficiency for 1% Gd in water
 - Gd → Big fat 8 MeV gamma signal when it decays, easy to detect and identify

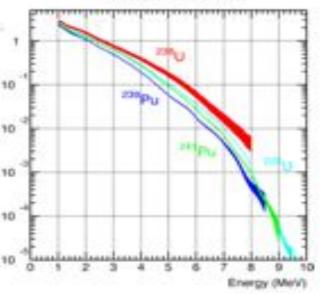
Illustrations in this section from presentations by K. Heeger, U. Wisc Y. Wang and J. Cao, ICHEP Beijing

Reactor Antineutrinos

Source

 $\overline{\mathbf{v}_{e}}$ from β -decays of n-rich fission products





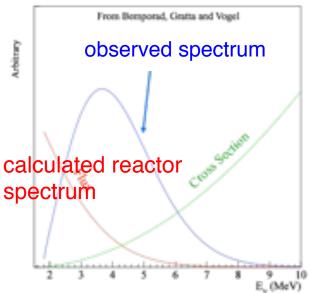
neutrinos/MeV/fission

pure ve source

> 99.9% of $\overline{v_e}$ are produced by fissions in 235 U, 238 U, 239 Pu, 241 Pu

Detection

inverse beta decay $\nabla_e + p \rightarrow e^+ + n$



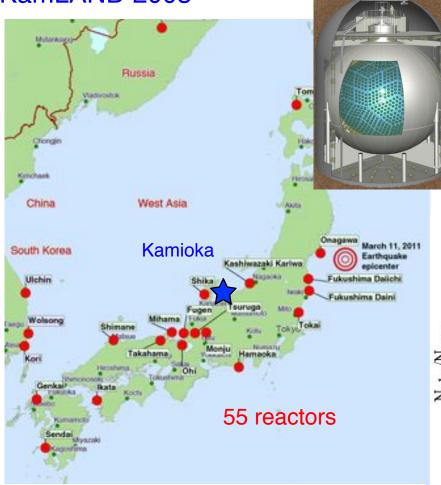
mean energy of $\overline{v_e}$: 3.6 MeV

only disappearance experiments possible

Observation of Reactor $\overline{v_e}$ Disappearance

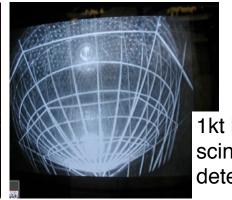


KamLAND 2003

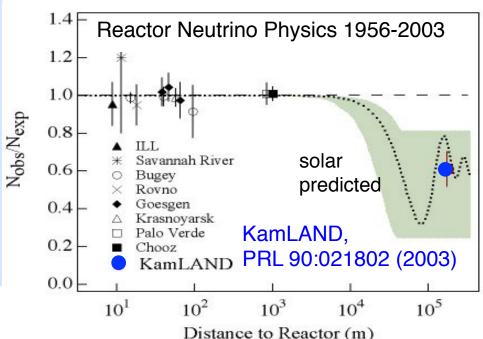


mean, flux-weighted reactor distance ~ 180km





1kt liquid scintillator detector



300 liters of liquid scintillator loaded with cadmium

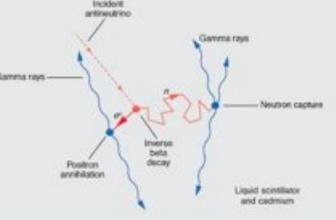
Long history: Hanford Experiment

(1953) F. Reines, C. Cowan

inverse beta decay

$$v_e + p \rightarrow n + e^+$$
 annihilates promptly
$$n + {}^{108}Cd \rightarrow {}^{109m}Cd \rightarrow {}^{109}Cd + v$$





signal: delayed coincidence between positron and neutron capture on cadmium

0.41+/- 0.20 events/minute

high background (S/N \sim 1/20) made the Hanford experiment inconclusive

The Savannah River (version 2 of Reines) Detector

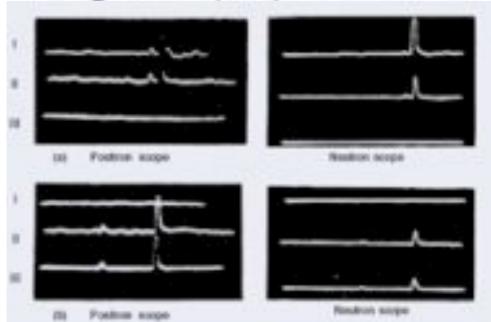
Improved design (1956), bigger reactor

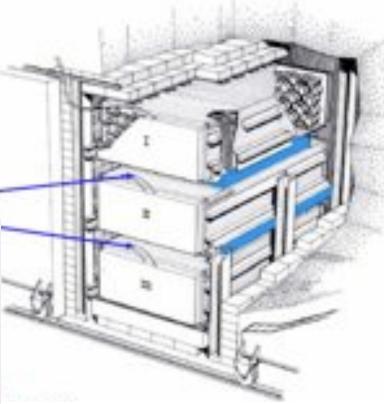
tanks I, II, and III were filled with liquid scintillator and instrumented with 5" PMTs

target tanks (blue) were filled with water+cadmium chloride

inverse beta decay

$$\overline{v_e} + p \rightarrow e^+ + n$$



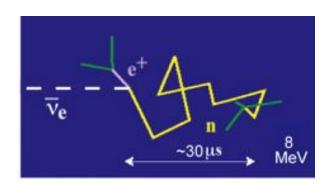


inverse beta decay would produce prompt and delayed signal in neighboring tanks

low E anti-neutrino detection: Inverse-β reaction in liquid scintillator

$$\overline{v}_{e} + p \rightarrow e^{+} + n$$

Illustrations in this section from presentations by Y. Wang, and J. Cao, ICHEP Beijing K. Heeger, U. Wisconsin



 $\tau \approx 180 \text{ or } 28 \,\mu\text{s}(0.1\% \text{ Gd})$

n + p
$$\rightarrow$$
 d + γ (2.2 MeV)
n + Gd \rightarrow Gd* + γ 's (8 MeV)

Neutrino Event: coincidence in time, space and energy

Observable V Spectrum

Neutrino energy:

$$E_{\overline{v}} \cong T_{e^+} + T_n + (M_n - M_p) + m_{e^+}$$

10-40 keV

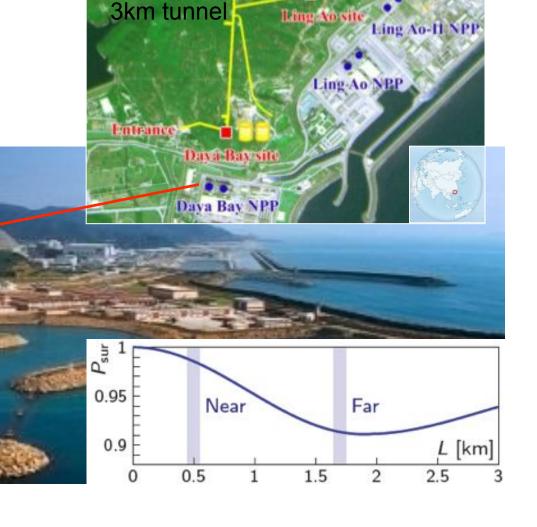
1.8 MeV: Threshold

Y. Wang, ICHEP Beijing

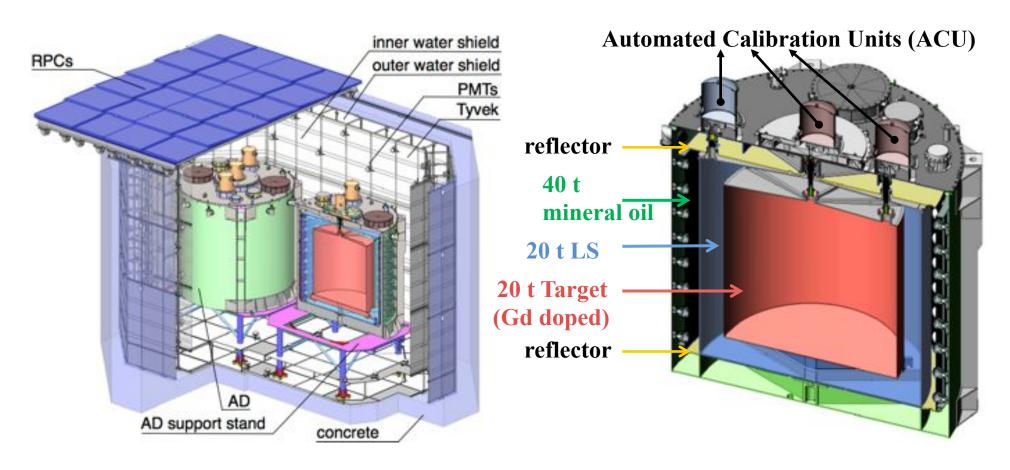
The Daya Bay Experiment

Far site

- 6 reactor cores, 17.4 GW_{th}
- Relative measurement
 - 2 near sites, 1 far site
- Multiple detector modules
- Good cosmic shielding
 - 250 m.w.e @ near sites
 - 860 m.w.e @ far site
- Redundancy



The Daya Bay Detectors



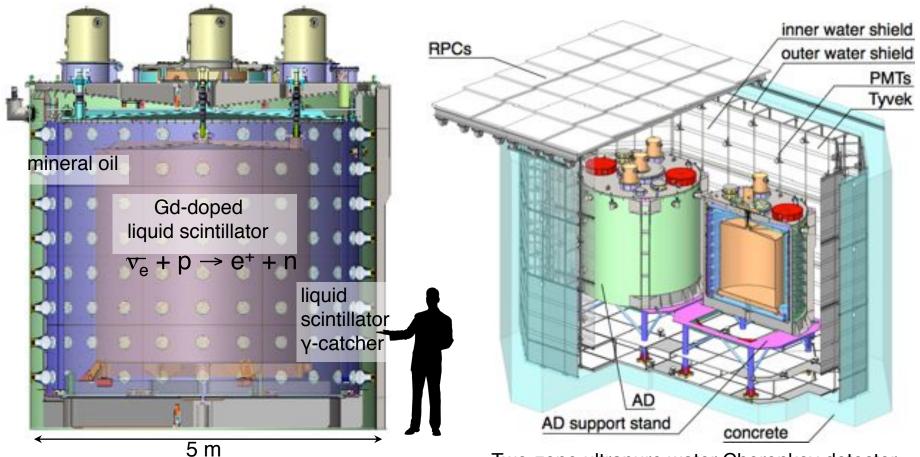
- Multiple AD modules at each site to check uncorr. syst. err.
 - Far: 4 modules, near: 2 modules
- Multiple muon detectors to reduce veto eff. uncertainties
 - Water Cherenkov: 2 layers
 - RPC: 4 layers at the top + telescopes

Daya Bay Detectors



6 "functionally identical" detectors
Gd-LS defines target volume, no position cut

Dual tagging systems: 2.5 meter water shield and RPCs



target mass: 20 ton per AD

photosensors: 192 8"-PMTs

energy resolution: $(7.5 / \sqrt{E} + 0.9)\%$

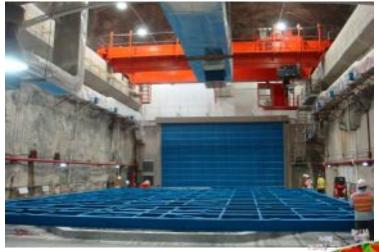
Two-zone ultrapure water Cherenkov detector multiple detectors allow comparison and cross-checks

Tunnel and Underground Lab.

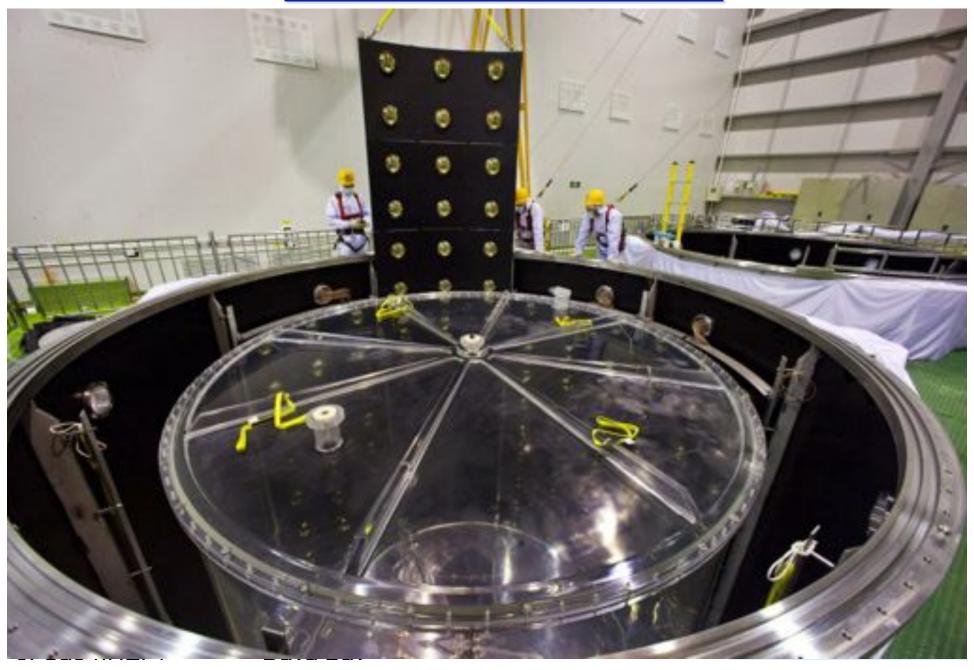






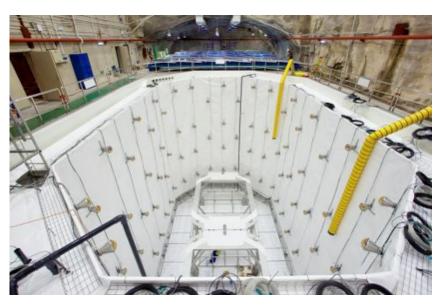


Antineutrino Detector



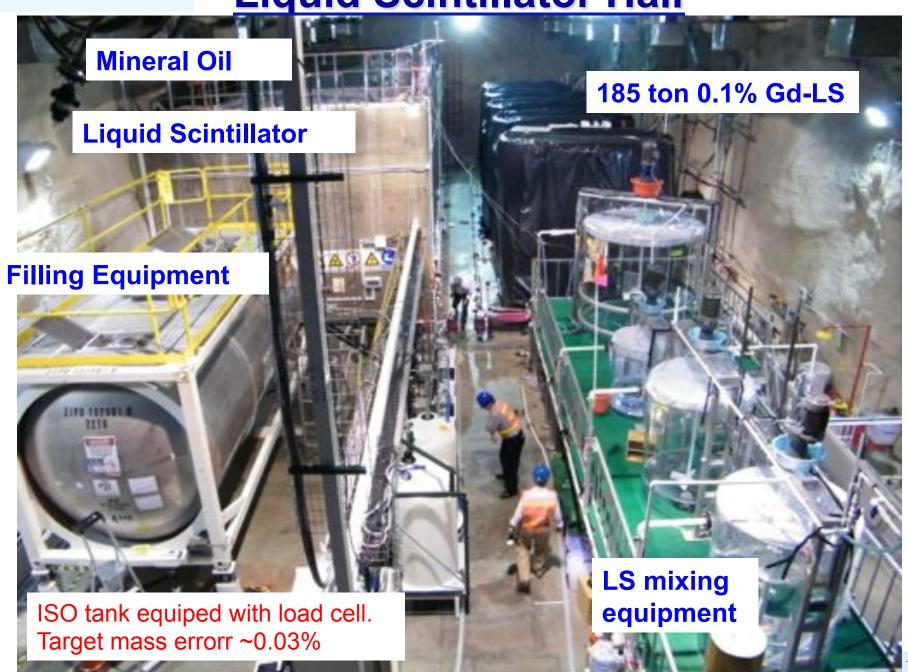
Muon System Installation





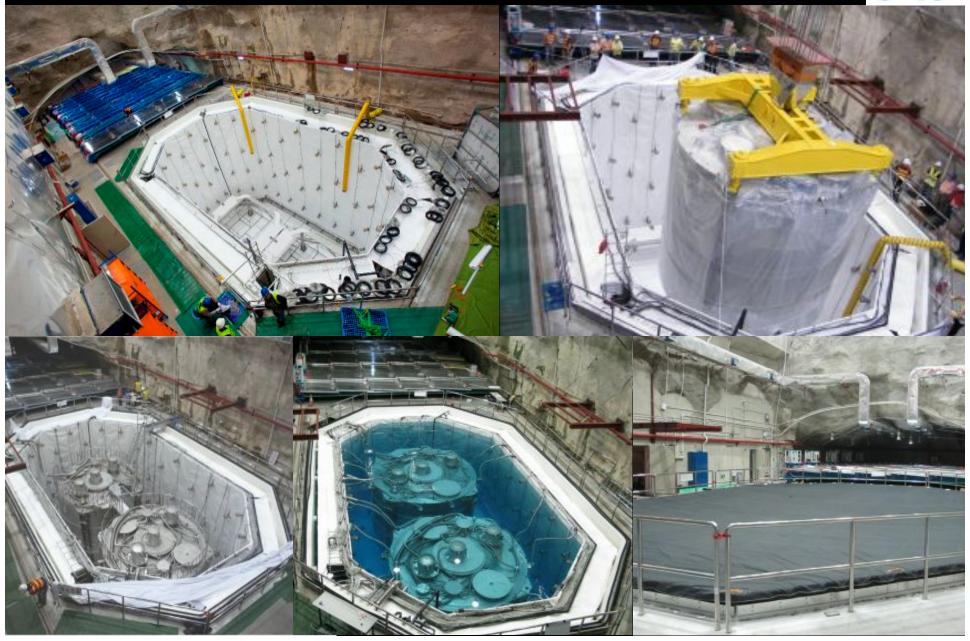


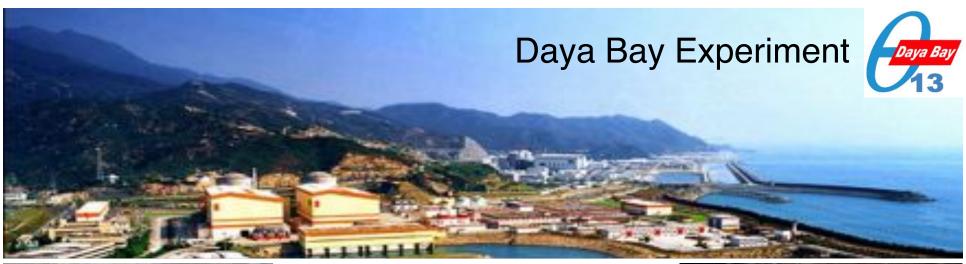
Liquid Scintillator Hall



Antineutrino Detector Installation - Near Hall

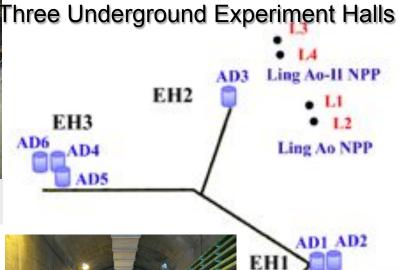








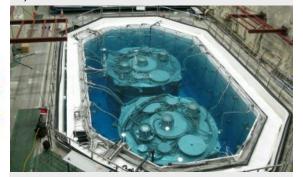
Hall 3: began 3 AD operation on Dec. 24, 2011







Hall 2: began 1 AD operation on Nov. 5, 2011



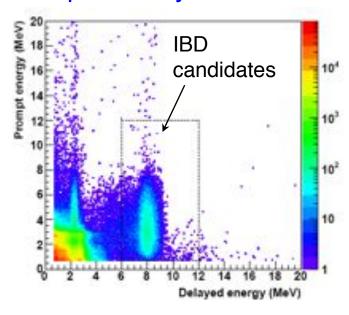
Hall 1: began 2 AD operation on Sep. 23, 2011

Daya Bay NPP

Antineutrino Candidates (Inverse Beta Decay)



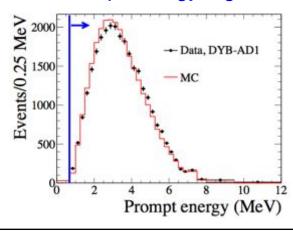
Prompt + Delayed Selection

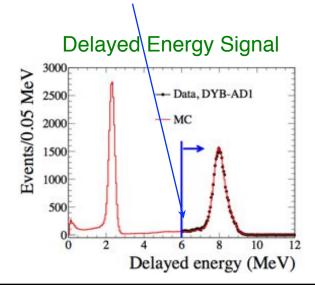


$$\overline{v_e} + p \rightarrow e^+ + n$$

Uncertainty in relative E_d efficiency (0.12%) between detectors is largest systematic.

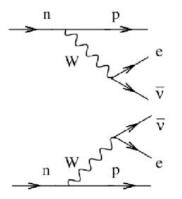
Prompt Energy Signal



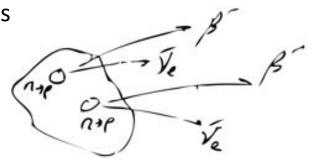


Double-beta decay, with and without neutrinos

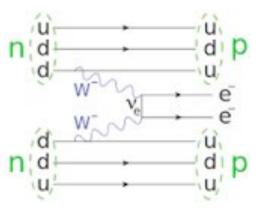
- Ordinary double-beta decay:
 - Simultaneous decays of 2 neutrons in a nucleus



 $2vββ - (A,Z) \rightarrow (A,Z+2)+2e^-+2\overline{v}_e$ SM Allowed and observed in select even-even isotopes

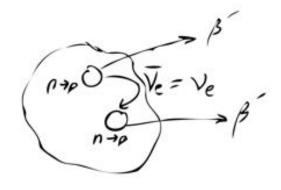


- Neutrinoless double-beta decay:
 - Antineutrinos annihilate (requires antinu=nu!)



$$0v\beta\beta - (A,Z)\rightarrow (A,Z+2)+2e^{-}$$

 $\Delta L = 2$



Illustrations in this section from presentations by

- J. Detwiler, UW
- C. M. O'Shaughnessy, U.N. Carolina, June 2015
- S. Elliott, LBNL

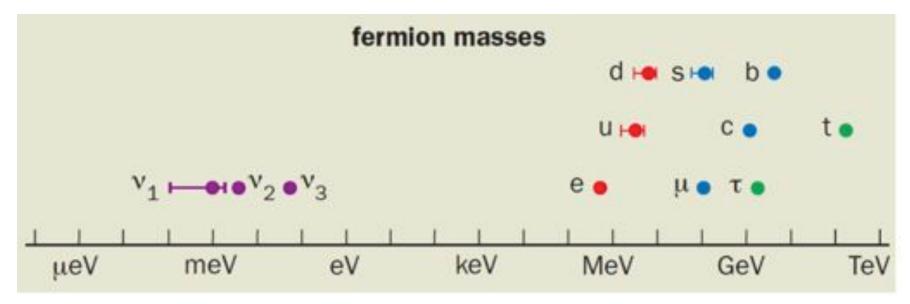
Majorana neutrinos

- 1937: Ettore Majorana proposed possibility of spin-1/2 particles (fermions) could be their own antiparticles
 - Neutral spin-1/2 particles are described by a real wave equation (Majorana equation)
 - So, identical to their antiparticle: wave functions of particle and antiparticle are complex conjugates
 - Creation and annihilation operators are identical
 - Standard Model assumes fermions are Dirac particles, with wavefunctions and creation/annihilation operators distinct
- All Standard-Model fermions except neutrinos are known to have distinct antiparticles
 - BUT: non-zero neutrino mass is non-SM
 - Hypothesized "sterile" neutrinos (not observed because they do not display weak interactions, only gravity) could be Majorana
 - In supersymmetry models, neutralinos (superpartners of gauge bosons and Higgs bosons) are Majorana.



Ettore Majorana
1906—1938 (?)
Majorana disappeared
from a ferryboat in
1938. Friends
presumed suicide.
Conspiracy theorists
believe he may have
hidden in Argentina or
Italy and lived until
1959.

Why are neutrino masses so much smaller than everybody else?

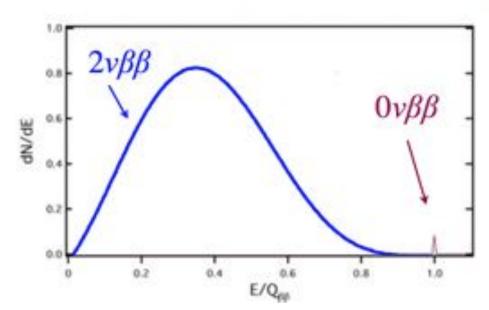


- Do they relate to Supersymmetry?
- Are they Majorana fermions?
- Could this be connected to matter/antimatter asymmetry in the Universe (do neutrinos exhibit CP violation)?

0νββ Discovery Considerations

- Need large, highly efficient source mass
- Desire extremely low (near-zero) backgrounds in the 0vββ peak region

- ⇒ Signal background 1:1 or better
 ⇒ Best possible resolution, ΔE, to minimize region of interest
- Want best possible energy resolution and/or kinematical method to discriminate 0vββ from 2vββ



Tonne scale experiments require backgrounds of ≤1 cts / ROI-t-y

> Need independent observations from different isotopes

One approach: Use ultra-pure Germanium detectors

Source = Detector

Intrinsically high purity, elemental Ge

Demonstrated ability to enrich to 86% ⁷⁶Ge

0.16% energy resolution at 2039 keV

Well-understood technologies

Commercial Ge diodes

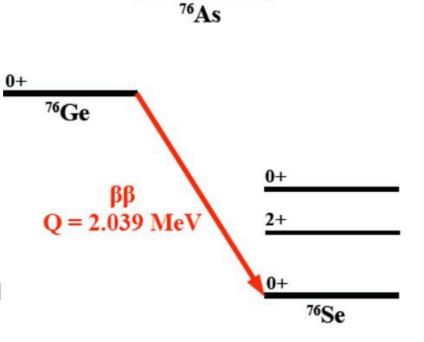
Majorana Experiment design:

Modules of 57 close-packed, I.I kg, segmented n-type HPGe detectors enriched to 86% ⁷⁶Ge

Independent cryostats made of ultra-clean electroformed Cu

Low background passive lead + electroformed Cu shield and 4π active veto

Located deep underground (4500-6000 mwe)



Deep underground site: Homestake Gold Mine → Sanford Underground Research Facility

Ross Campus

- Located in Lead, South Dakota
- Site for Ray Davis solar neutrino experiment
- Deepest facility in USA

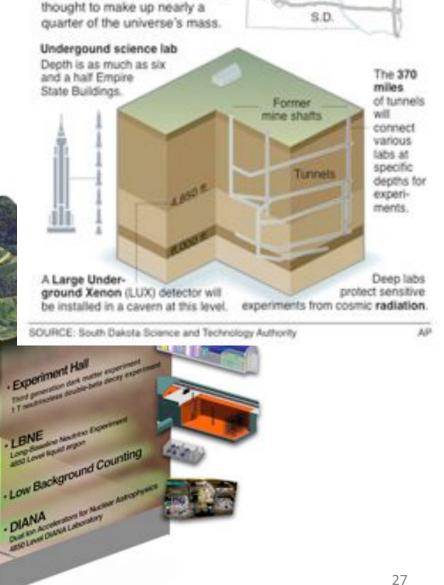
Davis Campus

autricies southe beta decay

argo Underground Xenon expe

ments in the Dak

11/24/15



From gold mine to science laboratory

Pieme

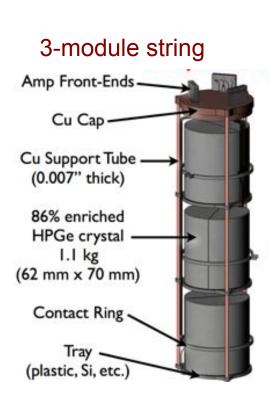
Work crews are restoring a

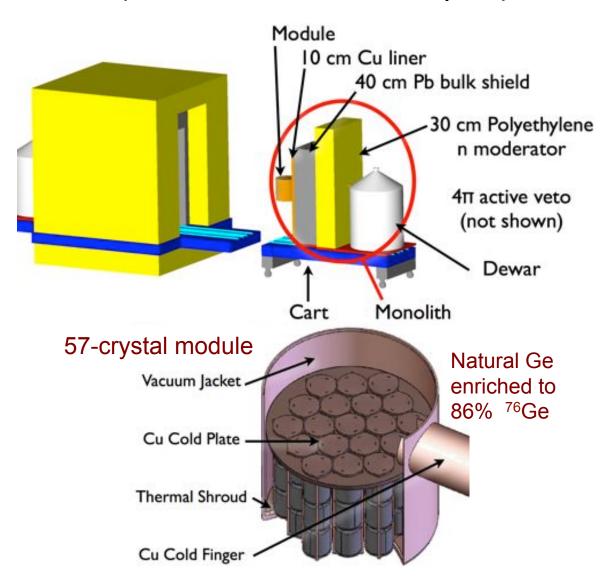
former gold mine in Lead, S.D., to house a detector to catch

what scientists call dark matter,

Majorana Experiment (US-Canada-Russia-Japan)

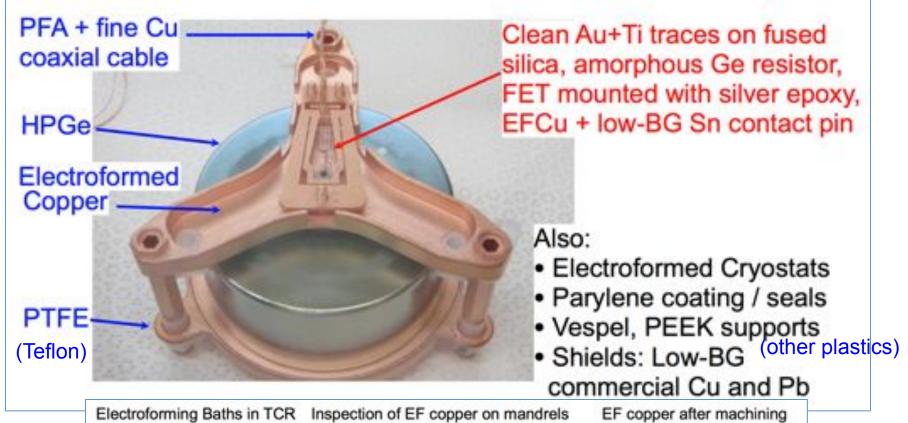
 Ge detectors in super-low BG site (Davis campus in SURF, 4850 ft deep)





Modular – expand to 1 ton total Ge

Individual Ge crystal in Cu mounting











- Th decay chain (ave) ≤ 0.1 µBg/kg
- U decay chain (ave) ≤ 0.1 µBq/kg

For now: demonstrator project with 44 kg Ge

Goals: - Demonstrate backgrounds low enough to justify building a tonne scale experiment.

- Establish feasibility to construct & field modular arrays of Ge detectors.
- Searches for additional physics beyond the standard model.
- Located underground at 4850' Sanford Underground Research Facility
- Background Goal in the 0vββ peak region of interest (4 keV at 2039 keV)
 3 counts/ROl/t/y (after analysis cuts) Assay U.L. currently ≤ 3.5
 scales to 1 count/ROl/t/y for a tonne experiment



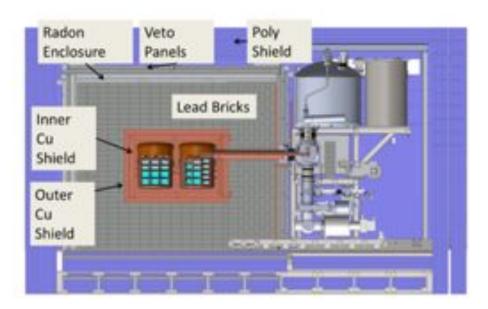
- 29 kg of 87% enriched 76Ge crystals
- 15 kg of ^{nat}Ge
- Detector Technology: P-type, point-contact.

· 2 independent cryostats

- ultra-clean, electroformed Cu
- 20 kg of detectors per cryostat
- naturally scalable

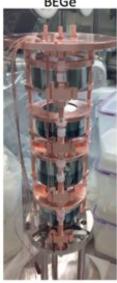
Compact Shield

 low-background passive Cu and Pb shield with active muon veto





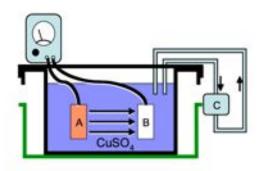
String with 3 EnrGe PPCs and 1 NatGe BEGe





Up to 5 detectors are mounted in 'strings'

Ultra-low-BG
high-purity
electroformed
Cu mountings
for Ge crystals



- Semiconductor-grade acids, recrystallized CuSO₄, high-purity copper stock
- Baths circulated with microfiltration, barium scavenge; cover gas
- Active plating manipulation, surface machining, cleaning, and passivation

- 232Th < I µBq/kg
- Recently improved bath chemistry: requires less surface finishing
- Improved starting stock quality and handling



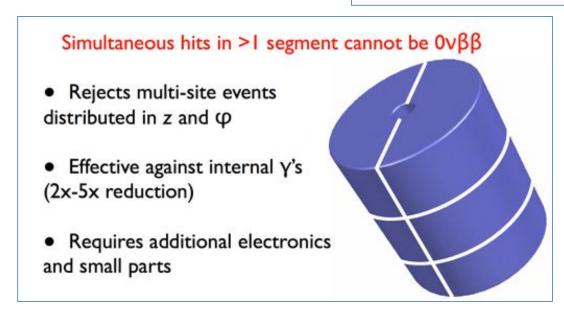
Event selection for background rejection

Main BGs:

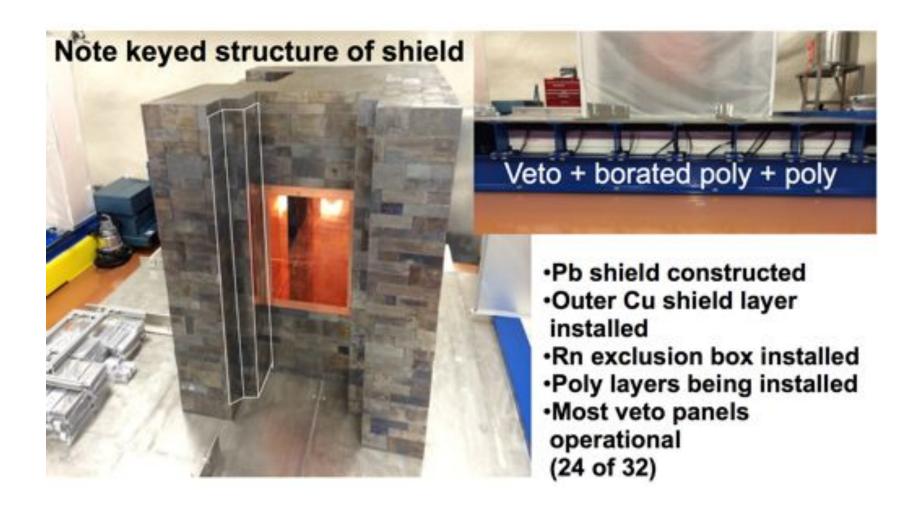
- High energy gammas
- Neutrinos
- Muon interactions
 - Spallation products

Simultaneous hits in > I detector cannot be 0vββ Effective for: • High energy external γ's, e.g. ²⁰⁸Tl and ²¹⁴Bi (2x-5x reduction) • Some neutrons • Muons (10x)

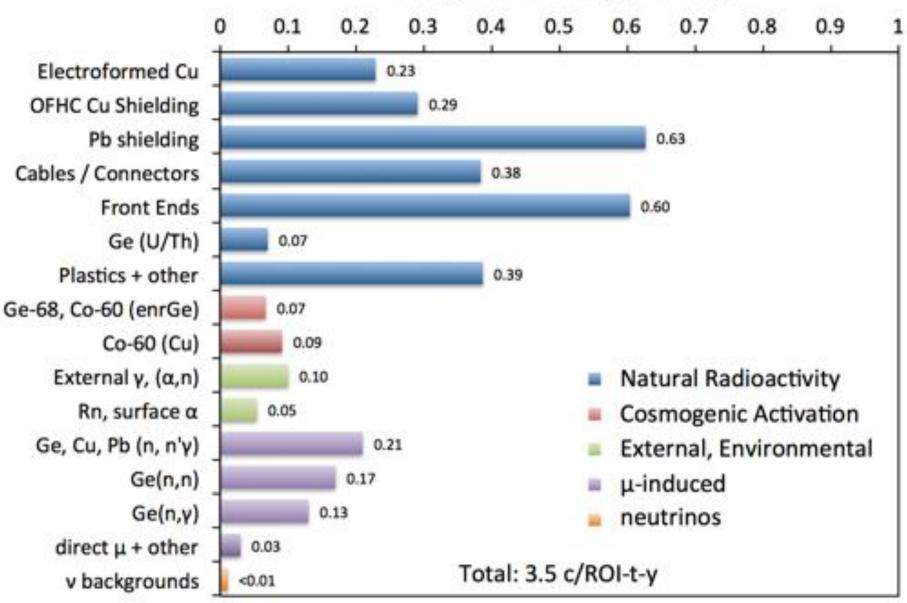
32



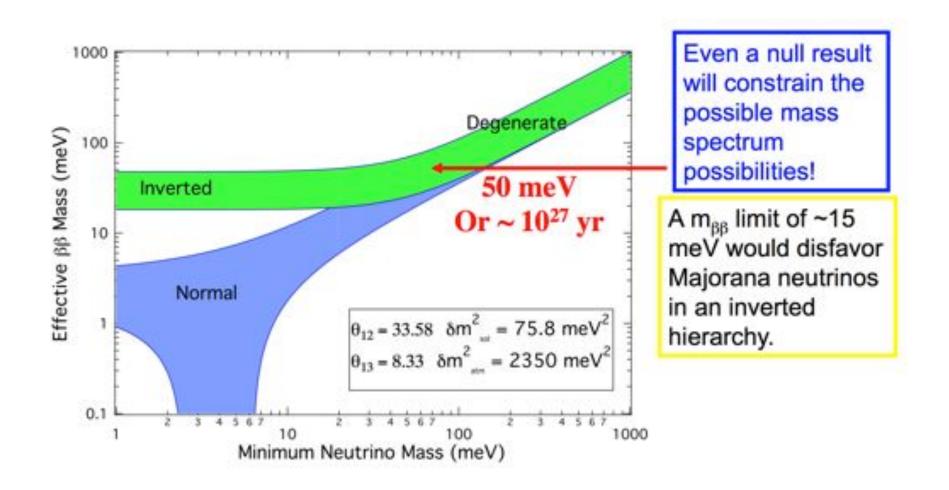
Shielding against neutrons and gammas



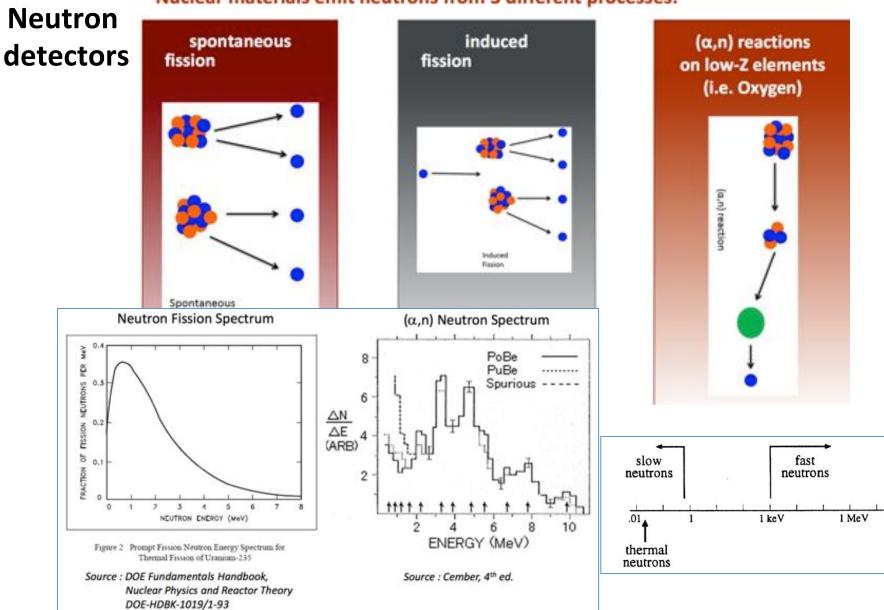




Majorana goals: milli-eV neutrino mass determination, as well as Majorana/Dirac determination



Nuclear materials emit neutrons from 3 different processes:



- Neutron detectors use nuclides that capture n and emit charged particles
- Reactions used include (σ 's below are for thermal n's)
 - − 10B(n, α) [σ ~4000 b] \rightarrow 2.31 and 2.79 MeV Q values \rightarrow ~1.5 MeV alphas
 - − 6Li(n,α) [σ ~1000 b] → 4.78 MeV Q → ~2 MeV alphas
 - − 3He(n,p) [σ ~5000 b] \rightarrow 0.76 MeV Q \rightarrow ~ 0.6 MeV protons
 - 157Gd (many reactions, to gammas and e's) [σ ~250,000 b!]
- Boron (BF₃ gas, or B-coated) proportional tubes
 - Inner surface of ionization chamber is covered with a thin coat of boron, or BF3 gas is used to fill the PC
 - Neutron is captured by a boron atom and an energetic alpha particle is emitted ${}_{5}^{10}B + {}_{0}^{1}n \rightarrow {}_{3}^{7}Li^{+++} + {}_{2}^{4}He^{++} + 5e^{-}$
 - Alpha particle and e's cause ionization within the chamber

www.centronic.co.uk

- ³He proportional tubes
 - Example: SNO neutrino detector
 - Neutrino Signal: Neutron from NC interaction
 - Neutrons capture via 3He(n,p)3H in the NCD and produce 573 keV p + 191 keV tritium ionization tracks.

Distinguishable Backgrounds:

• Tritium in 3He

3H decays deposit on average 6 keV in the gas but pile-up can produce proportional counter signals above threshold. Low-temperature purification of the 3He has resulted in negligible background levels.

Surface and Bulk Alpha Activity

232Th and 238U chains in the NCD walls, along with 210Po surface activity, produce α 's that underlie the neutron capture peak. These events can be rejected by event by event analysis of digitized pulses. (see "Event Identification by Pulse Shape Analysis")

Electrons and Gammas

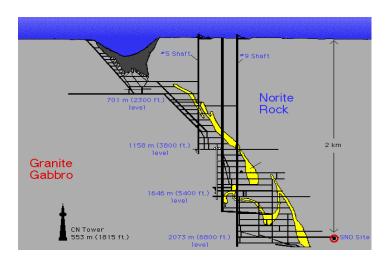
betas and gammas from the 232Th and 238U chains only deposit 764 keV through extensive multiple scattering. Less than 2x10-4 fall into the neutron window.

γ, Eγ=0.7-2.0 MeV

 ν or γ , E>2.22 MeV

www.sno.phy.queensu.ca/

Sudbury Neutrino Observatory



1000 tonnes D₂O

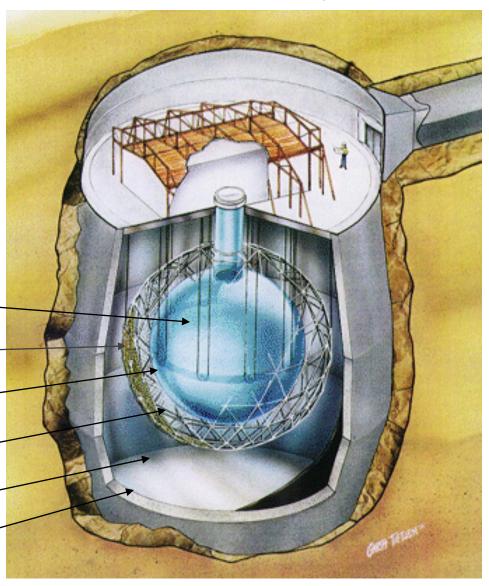
Support Structure for 9500 PMTs, 60% coverage

12 m Diameter Acrylic Vessel

1700 tonnes Inner Shielding H_2O

5300 tonnes Outer Shield H₂O

Urylon Liner and Radon Seal



SNO Neutral Current Detection Array

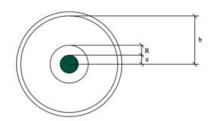


³He proportional counters detect the neutrons liberated from deuterium by neutrinos.

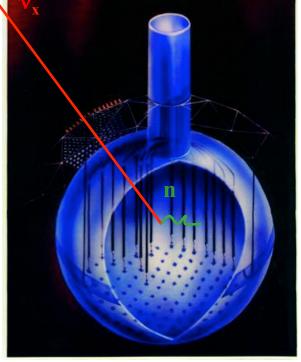
Total Length: 398 m

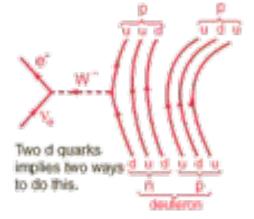
Vertical Strings: 40

n capture efficiency: $\varepsilon_n \sim 21\%$

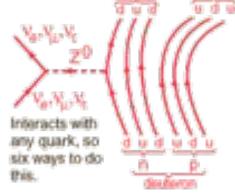








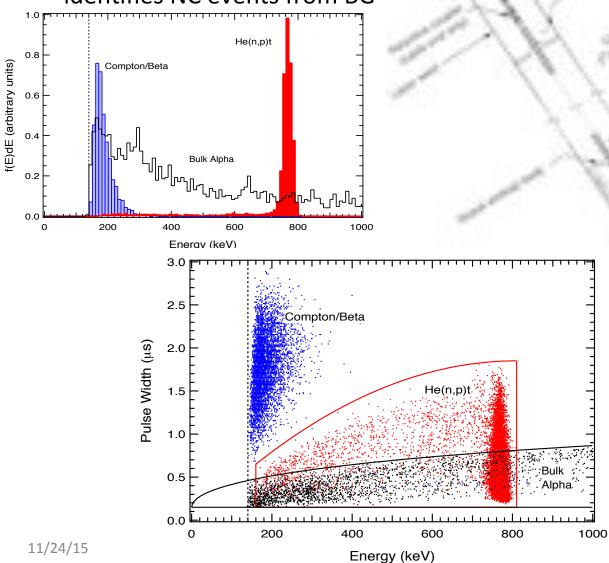
Charged current reaction, electron neutrinos only.



Neutral current reaction, all neutrinos. Elastic scattering with any neutrino. Neutral-current process (Z⁰ boson) is equally sensitive to all flavors of nu

SNO NC detectors

 Correlated E and pulse width identifies NC events from BG



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Neutron Detectors for Security Monitors

"SIGNIFICANT QUANTITIES" IAEA

"Significant Nuclear Material (SNM)" = Plutonium and highly enriched Uranium

SNM	Mass	Emissions from sphere	Neutron Flux @ 3 m
Pu	8 kg	2 x 10 ⁶ n/s	2 n/(s·cm²)
HEU	25 kg	100 n/s	10 ⁻⁴ n /(s·cm²)

Neutrons are not emitted with significant intensity from natural sources

[Source: Kouzes et al, doi:10.1016/j.nima.2007.10.026]

Spontaneous fission

- Even numbered plutonium isotopes
- 238U (very low intensity)

spontaneous fission isotope	neutrons/g-s
²³⁸ U	0.011
²³⁸ Pu	2,500
²⁴⁰ Pu	1,020
²⁴² Pu	1,700
²⁴⁴ Cm	11,000,000

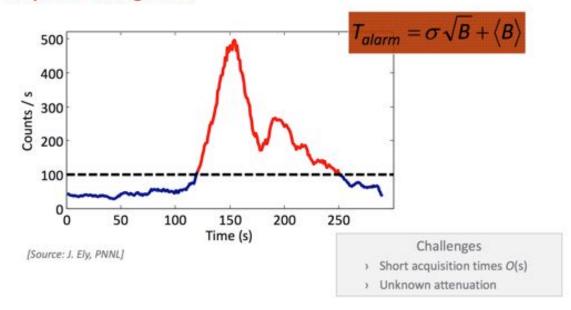
[Source: Nuclear Safeguards, Security, and Nonproliferation, JE Doyle]

• Why not use widely available, highly efficient gamma-ray detectors?

	Gamma-rays	Neutrons
Sources	Essentially all radioactive materials, even beta emitters yield bremsstrahlung Can detect trace quantities of material, e.g. fission products	Restricted to transuranic actinides • plutonium • californium
Background	Originates from radionuclides in the environmental	Originates from cosmic-ray spallation events
	Highly variable over small distances	Highly variable over altitude, to a lesser extent latitude
	Dependent on local terra & anthropogenic structures	Dependent on amount of surrounding high-Z material (ship effect)
	Weather variability due to radon	A. House and the second
Radiation Clutter	A serious problem	Manageable
Attenuation	Most effective with high-Z materials	Most effective with hydrogenous materials
	Typical attenuation lengths are O(cm)	Typical attenuation lengths are O(10 cm)

- Gamma spectroscopy (counting + precise energy spectrum) is useful when
 - Isotopic composition is measurement desired, or
 - Broad range of materials may be of interest, or
 - Source may be shielded by water or hydrogenous materials
- Neutron detectors (counting) are better when
 - Material mass is quantity of interest
 - Plutonium is source
 - Source may be shielded by high-Z materials

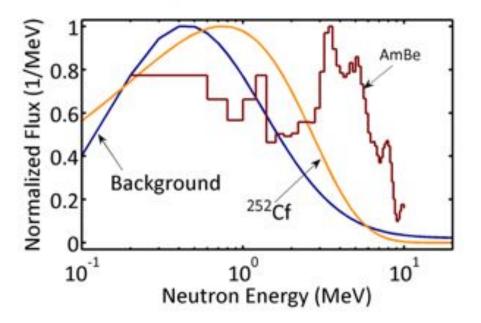
Goal: detect presence of a elevated neutron emissions with respect to background



Neutron Detectors for Security Monitors

It sounds easy but ...

- acquisition times are limited
 - what constitutes an alarm?
- backgrounds change during continuous acquisition
- some gamma-ray sources are huge
- innocent sources exist in commerce, albeit rarely



3He Neutron Detectors for Homeland Security Radiation Portal Monitors

Contemporary counters rely on ³He for thermal neutron detection

- hydrogenous moderator surrounds sample
- tubes embedded in moderator



- Current demands for ³He are greater than supply
 - only current source from tritium decay
 - tritium stocks dwindling
 - recent supply has been ~60 kL/year from stockpile

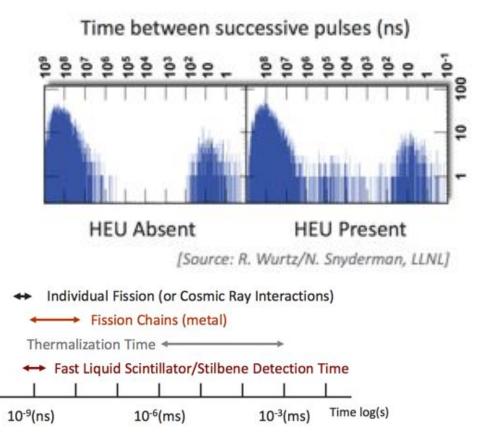
National security demand (was) rising

■ GE-Reuter Stokes projects ~65 kL/year

- Liquid scintillator detectors have much faster response but low efficiency compared to 3He
 - For high-intensity sources (>10⁴ n/sec not a problem
 - Can use correlations (multi-neutron emission) to identify fission sources: n counts are not Poisson-distributed: larger variance



[Source: L. Nakae, LLNL]

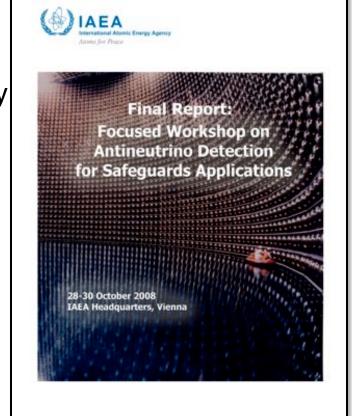


IAEA and Nuclear Non-Proliferation



IAEA Interest:

- Improved knowledge of input plutonium mass at reprocessing facility or repository
 - currently no better than 5-10%
- Research reactor power monitoring
 - currently uses intrusive tech.
- Verification of bilateral agreements
 - maybe future role for agency
- Detection with minimal overburden
 - allows widespread deployment







AGENDA

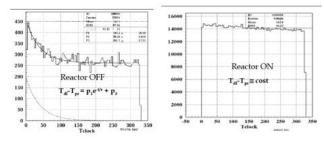
Ad Hoc Working Group on Safeguards Applications of Antineutrino Detectors, 14 September 2011, Vienna, Austria

Examples of monitoring detectors under development

CORMORAD: Gd/Plastic segments, no shield/veto

PANDA36: Gd/Plastic segments, no shield/veto

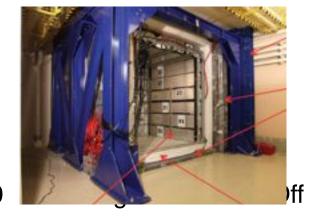
Nucifer: GdLS, lead/poly/ veto, 7m from RR



Rnd. Bkg x10² Rx On/Off Exp. signal/Obs.bkg ~1/100



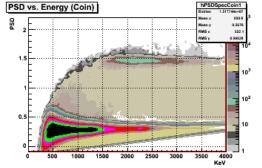
SNL/LLNL: LiZnS/Plastic, poly shield/veto



LLNL/SNL: Gd-Water Cerenkov, poly shield/veto



Exp.signal/Obs.bkg ~1/100



Neutron capture PSD gives >10² bkg rejection

Careful assessment of cosmogenic and reactor background required
Detectors with high selectivity of e+ and/or neutron capture required

Ref: Bowden, LLNL