

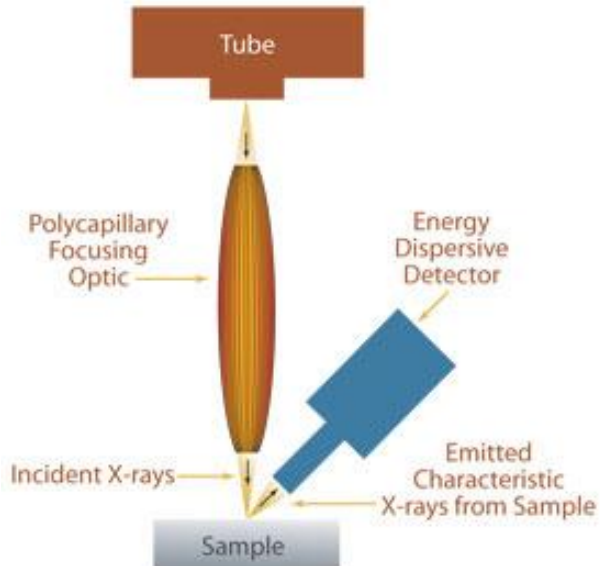


# Silicon drift detector response function to hard x-rays (with an introduction on quantitative MicroXRF And the Planetary Instrument for X-ray Lithochemistry)

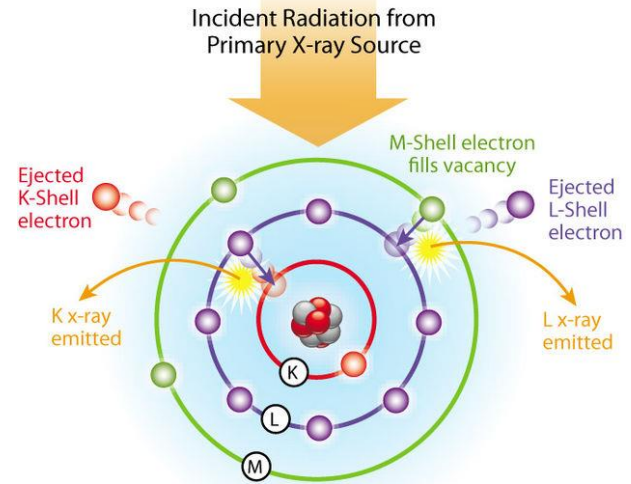
Nicolas Michel-Hart

with special thanks to Tim Elam for project guidance

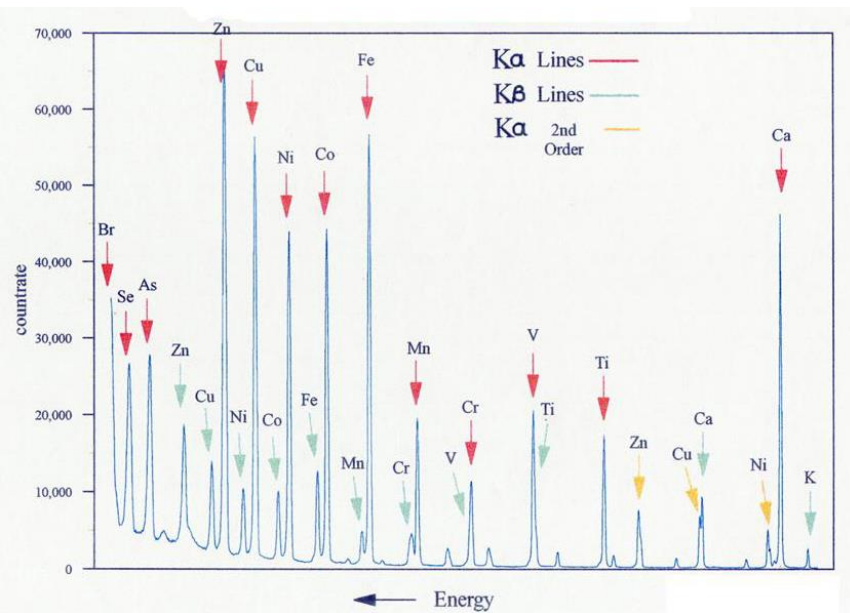
# X-ray Fluorescence and $\mu$ XRF



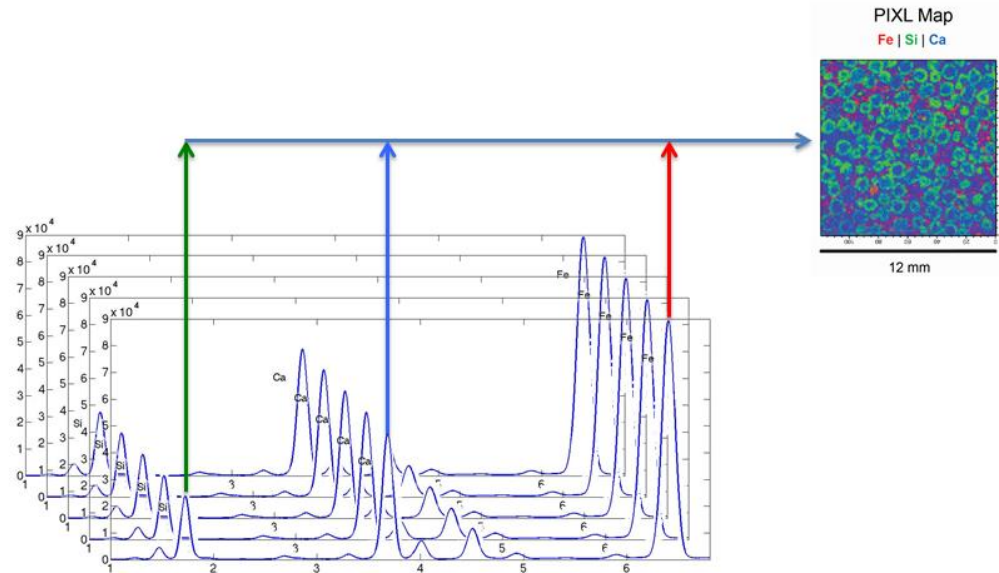
<https://xos.com/technologies/xrf/>



<http://in.niton.com/en/>

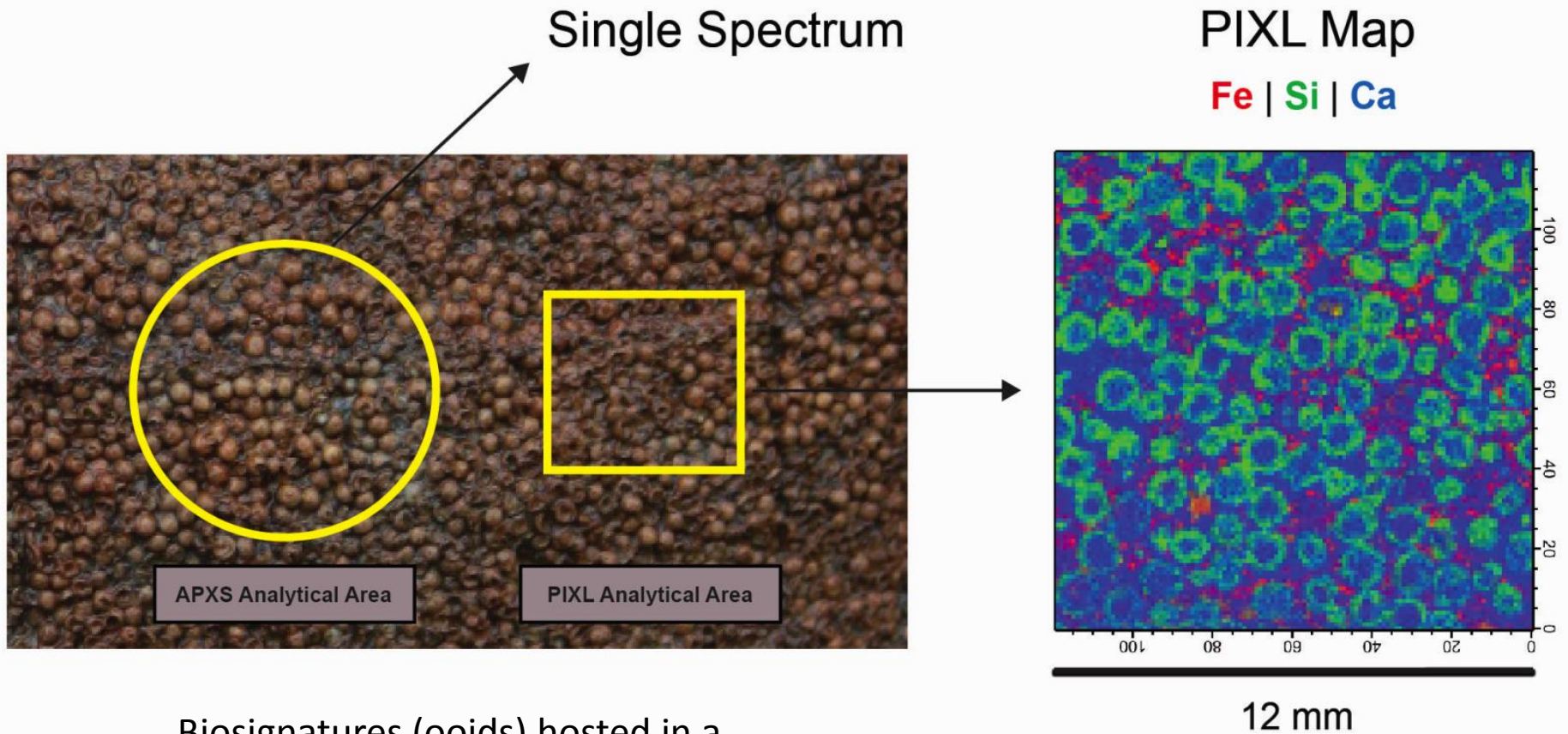


[https://en.wikipedia.org/wiki/X-ray\\_fluorescence#/media/File:XRFscan.jpg](https://en.wikipedia.org/wiki/X-ray_fluorescence#/media/File:XRFscan.jpg)



Elam, PIXL Seminar, Nov 2015

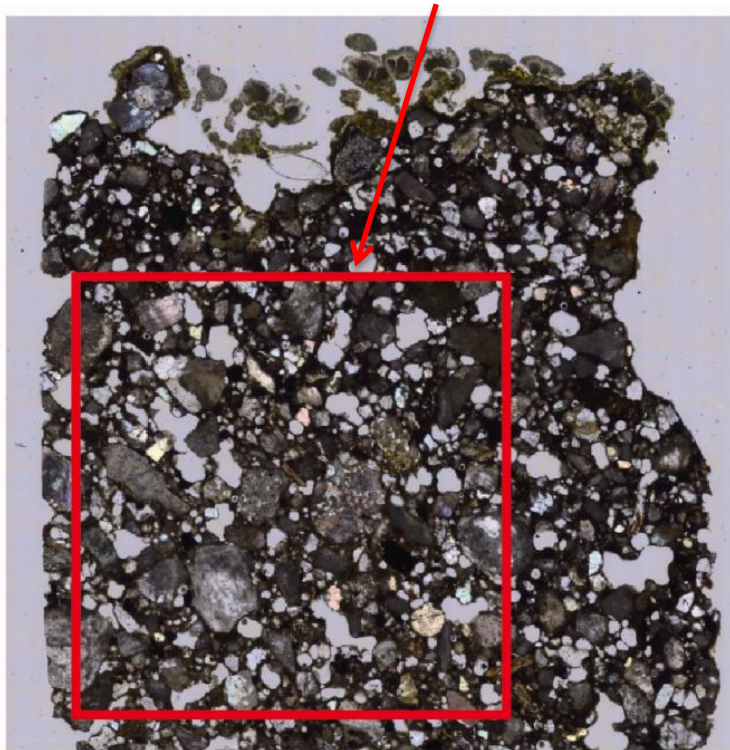
# Why Micro XRF?



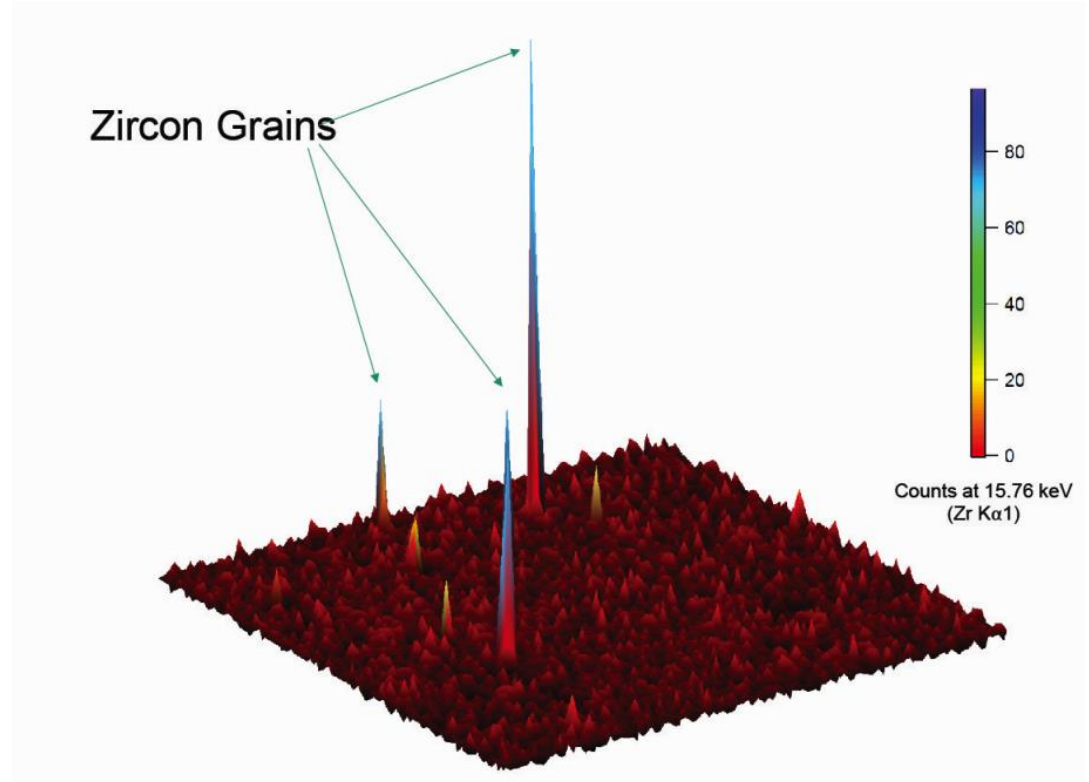
Biosignatures (ooids) hosted in a  
2.7 billion-year-old carbonate rock

# Some pixels can be very different

Area of element maps



Visible light image

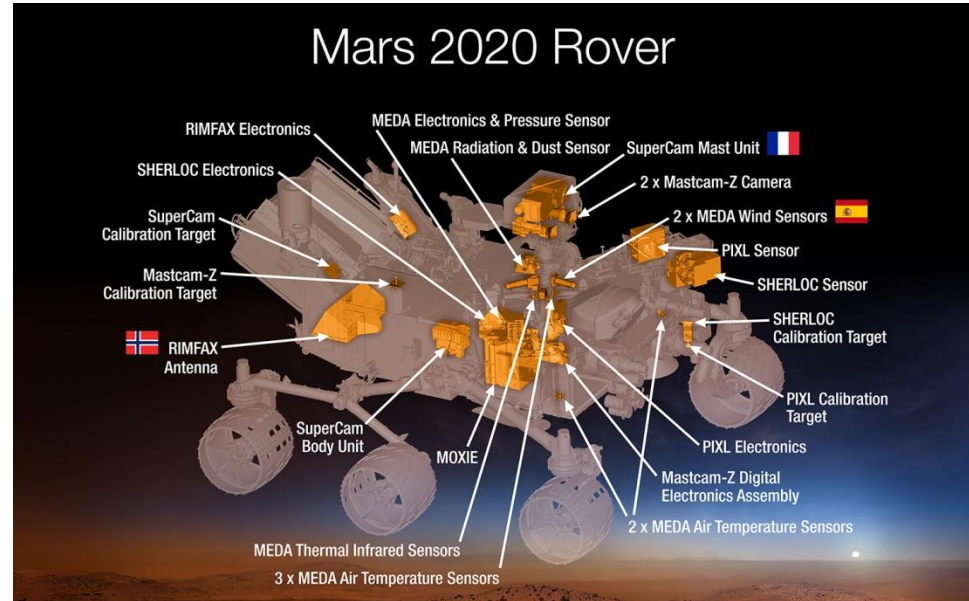


Recent microbialite from the Death Valley Area

# Mars 2020 Mission

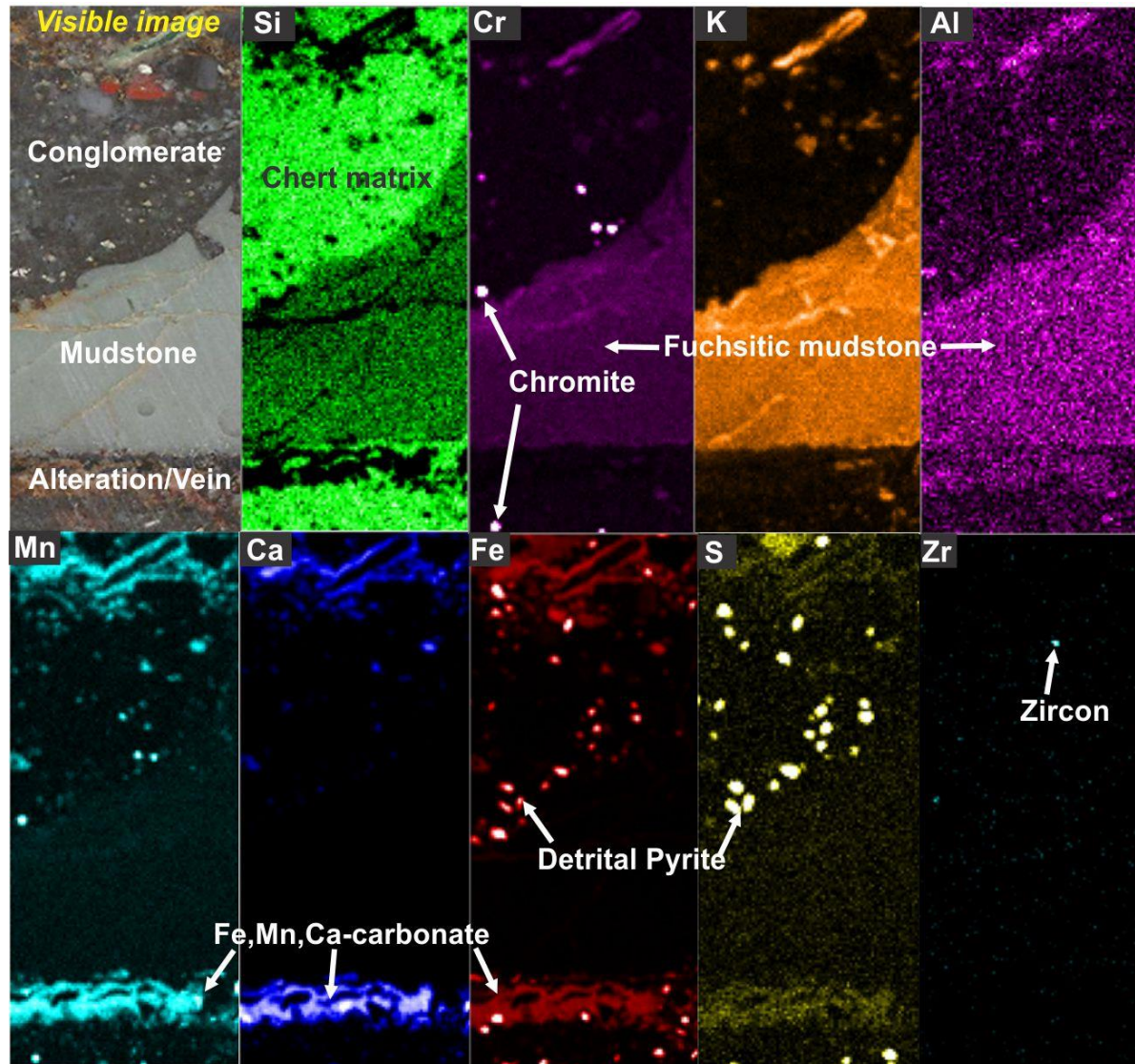
- M2020 Science Objectives

- **Habitability:** Characterize the geologic record for astrobiologically relevant environments and geologic diversity
- **Biosignatures:** Search for materials with high biosignature preservation potential
- **Sample Caching:** Obtain a pristine set of geologically diverse samples and cache for future return to Earth
- **Prepare for Humans:** Demonstrate *in situ* resource utilization technologies and characterize dust size and morphology



- Mission life: 1.5 Mars years/1005 Martian days
- Flight Instruments delivered by Fall 2018, Launch July 2020, Land February 2021
- Instrument Complement:
  - Mastcam-Z and Supercam for panoramic/stereo imaging and chemical analysis
  - MEDA for weather
  - RIMFAX ground penetrating radar
  - MOXIE technology experiment to produce Oxygen from CO<sub>2</sub>
  - SHERLOC and WATSON for UV Raman and high resolution imaging
  - **PIXL: Topic of today's talk and the coolest instrument on the Mars 2020 Mission!**

# Assessing Past Environments

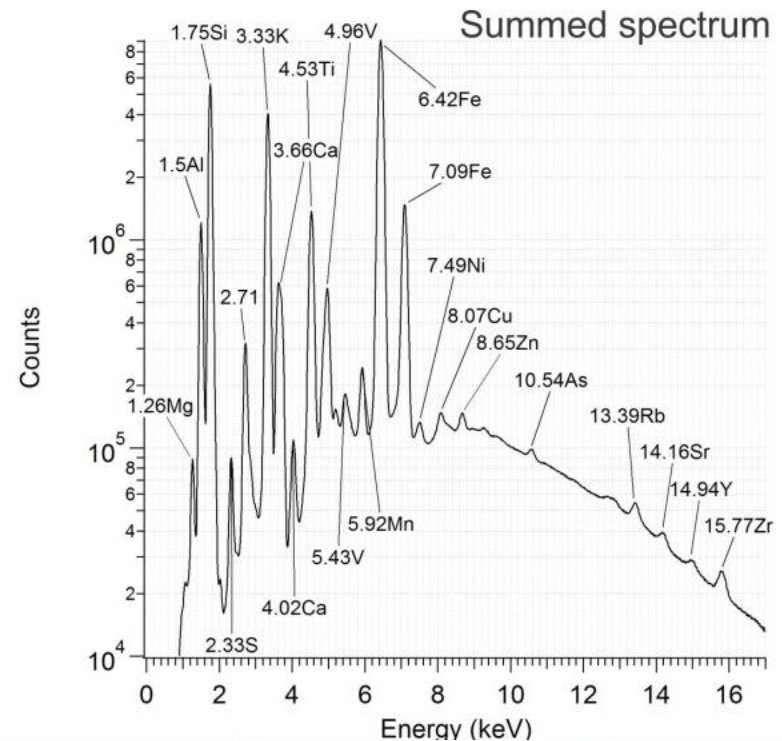
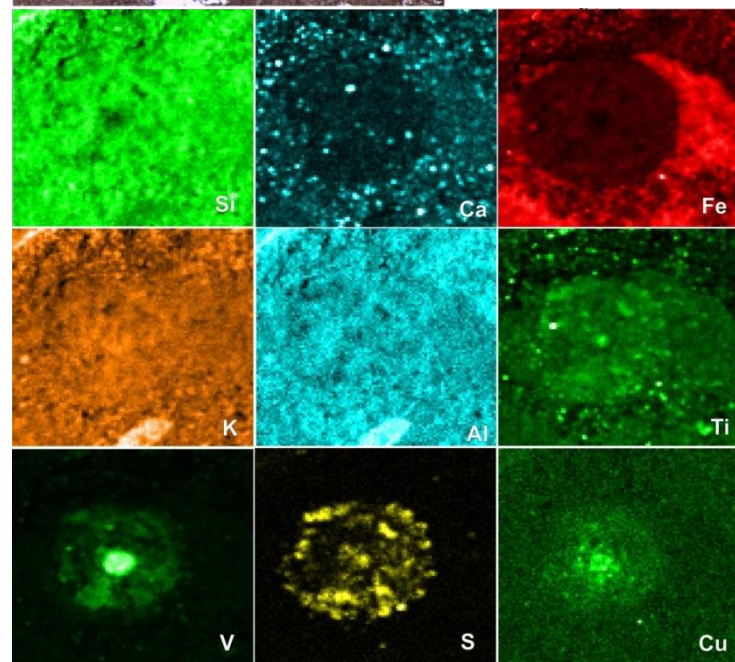
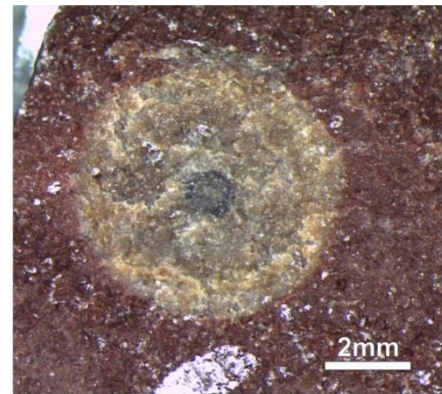


3.45 billion yr old  
conglomerate deposited on  
an ancient paleosol

Map size = 2 x 1 cm  
Step size = 150  $\mu$ m

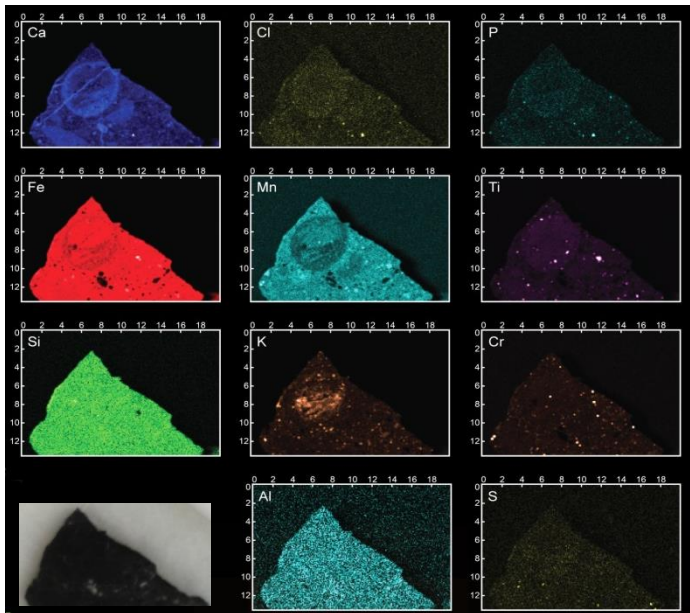
# Detection of Potential Chemical Biosignatures

Element maps reveal concentrated vanadium and copper in the black spot – a potential biosignature in sandstone

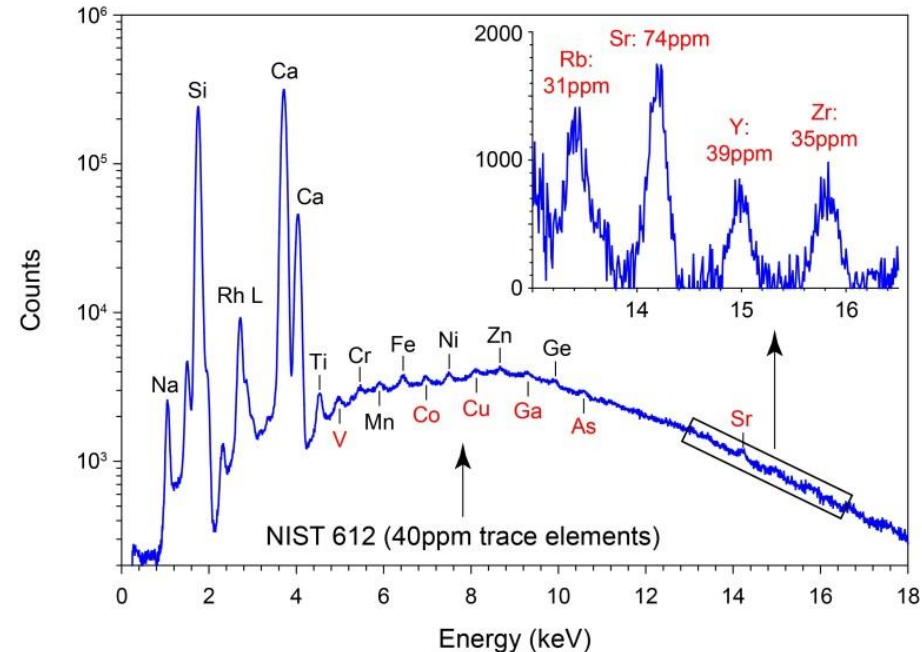


# PIXL Science, Data Analysis, and Hardware Team (partial)

- Abigail Allwood PI
- Joel Horowitz DPI
- Benton Clark Co-I
- Tim Elam Co-I
- John Grotzinger Co-I
- Robert Hodyss Co-I
- John Jorgensen Co-I
- Scott McLennan Co-I
- Michael Tice Co-I
- Allan Treiman Co-I
- David Flannery Co-I
- Yang Liu PS
- Marc Foote IPM
- Larry Wade ISE
- Douglas Dawson XRS Cog-E
- Moxtek X-ray tube
- XOS Polycapillary X-ray Optics
- Amptek Detectors
- Steve Battel HVPS
- Eric Hertzberg HVPS
- University of Michigan HVPS
- Danish Technical University Camera System



Element maps of Martian meteorite NWA 7034



PIXL breadboard NIST 612 spectrum

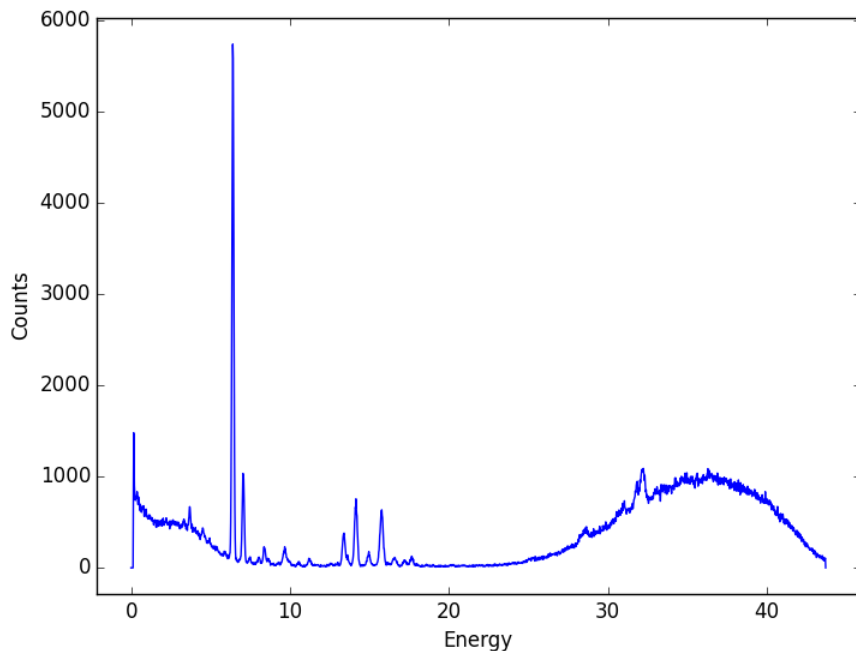
Elam, PIXL Seminar, Nov 2015



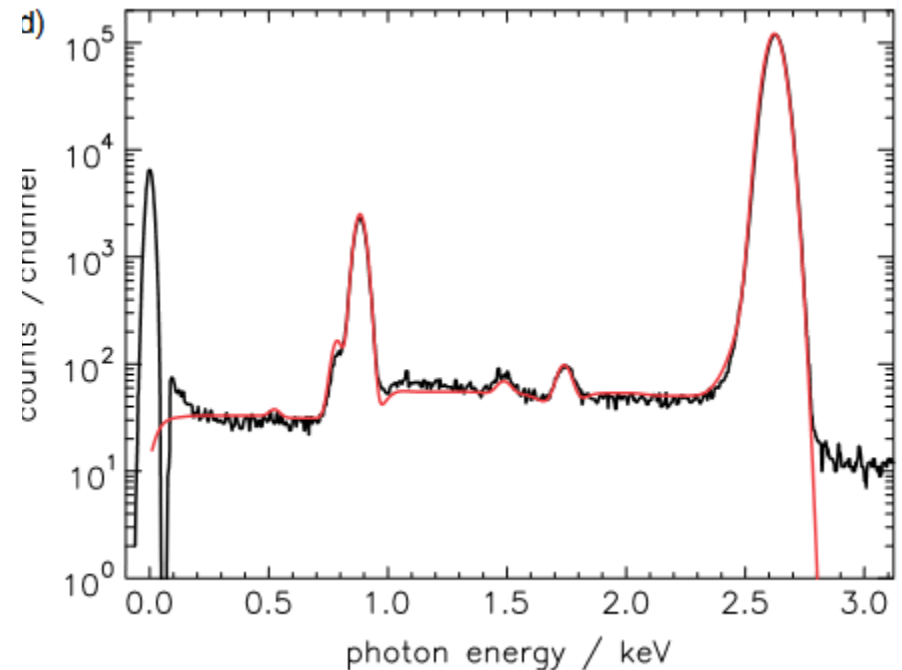
# The Problem

- Significant low energy counts when using high energy sources is not understood
- Differences observed in soft and hard x-ray output functions using the same detectors
- SDD response functions well researched and modeled for low energy x-rays
- Not so for high energy x-rays

Doped Opalinus Clay



W. T. Elam, unpublished raw data, 2015



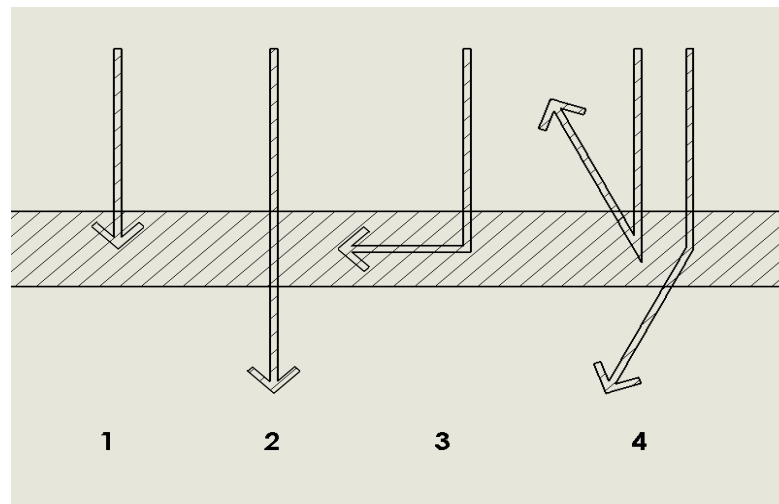
F. Scholze, M. Procop, X-Ray Spectrometry. 2009, 38, 312-321

# Hypothesis

**Any photon incident to the detector will have one of four fates:**

1. Photon is fully absorbed, its full energy is captured by the detector
2. Photon passes through the detector with no interaction, no energy is captured by the detector
3. Photon scatters inelastically in the detector, the scattered photon is then absorbed without escaping the detector. All of the incident photon energy is captured by the detector and measured as one count.
4. Photon scatters inelastically in the detector; the scattered photon then escapes the detector with no subsequent interactions. Only the energy transferred to an electron during the scattering event is captured and measured by the detector.

**We think case #4 is the cause of low energy counts being registered when there are no low energy photons present.**



# The physics of the four processes

## Input photon beam absorption

$$N_{absorbed} = N_0(1 - e^{(-\sigma\rho L)})$$

## Photon Compton scattering

$$\frac{d\sigma_c^Z}{d\phi} = \pi r_e^2 S(x, Z) \sin \phi \left[ (1 + \alpha(1 - \cos \phi))^{-2} \left[ 1 + \cos^2 \phi + \frac{\alpha^2(1 - \cos \phi)^2}{1 + \alpha(1 - \cos \phi)} \right] \right]$$

$$\frac{dN}{d\phi} = N \left( 1 - e^{-\frac{d\sigma_c^Z}{d\phi} \rho_{atom} L} \right)$$

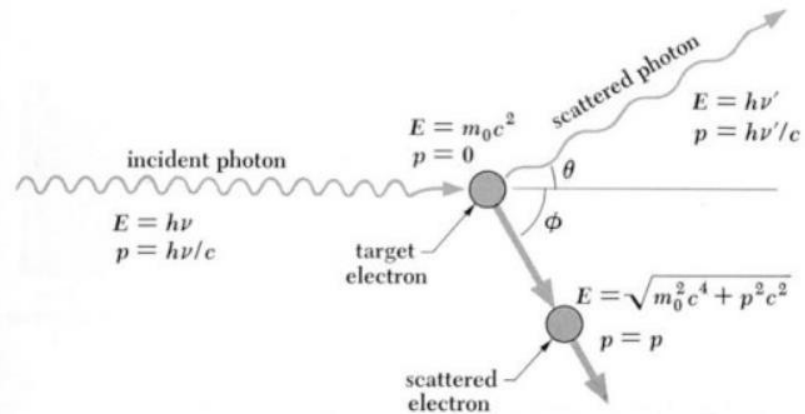
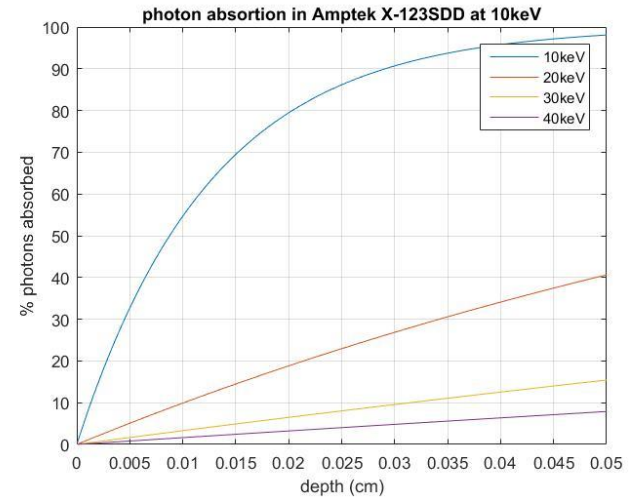
## Scattered photon and electron energy

$$E_{scattered} = E_0 \frac{1}{1 + \alpha(1 - \cos \phi)}$$

$$E_{electron} = E_0 \frac{\alpha(1 - \cos \phi)}{1 + \alpha(1 - \cos \phi)}$$

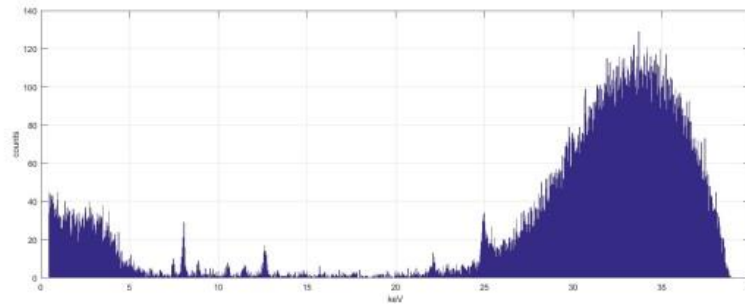
## Scattered photon escape

$$N_{escape} = N_{scattered} e^{(-\sigma_{escape\_photo} \rho EL)}$$



# Numerical modeling of the problem

- We don't have access to a monochromatic source
- Must work backwards from the experimental source peak to derive the incident x-ray beam

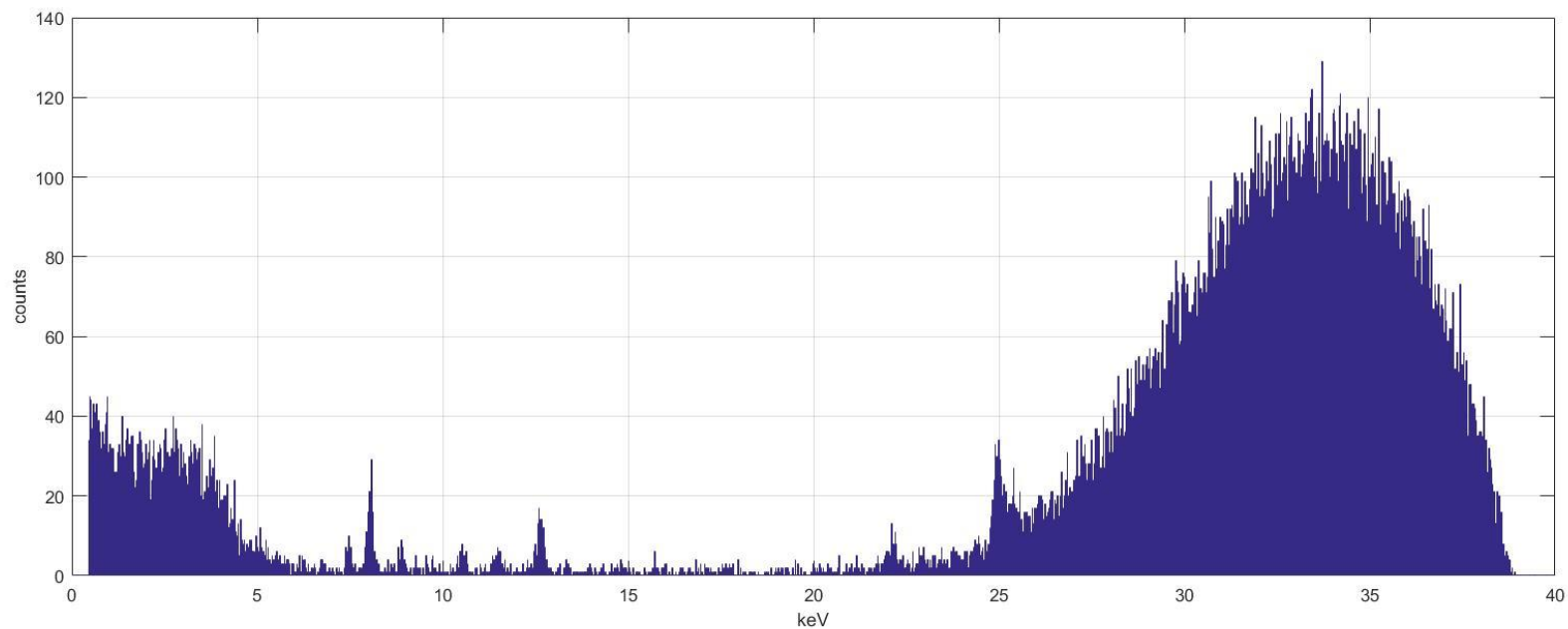
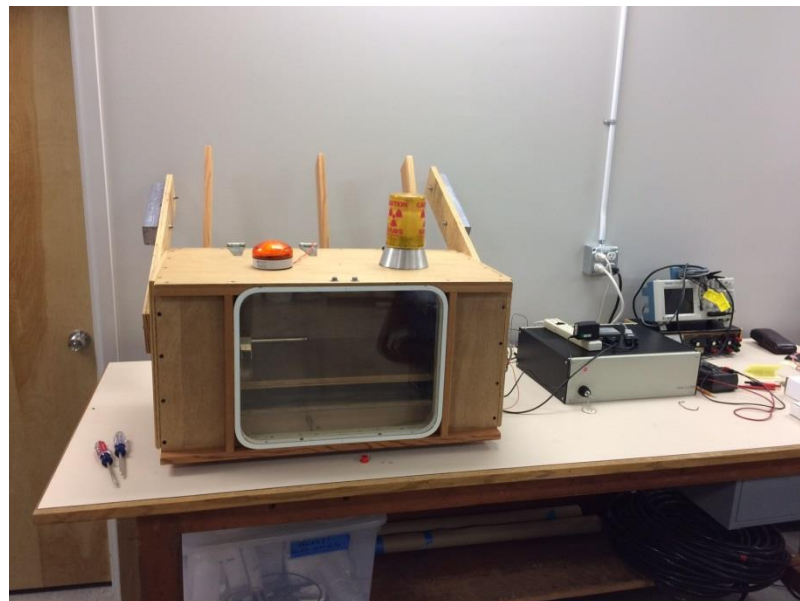
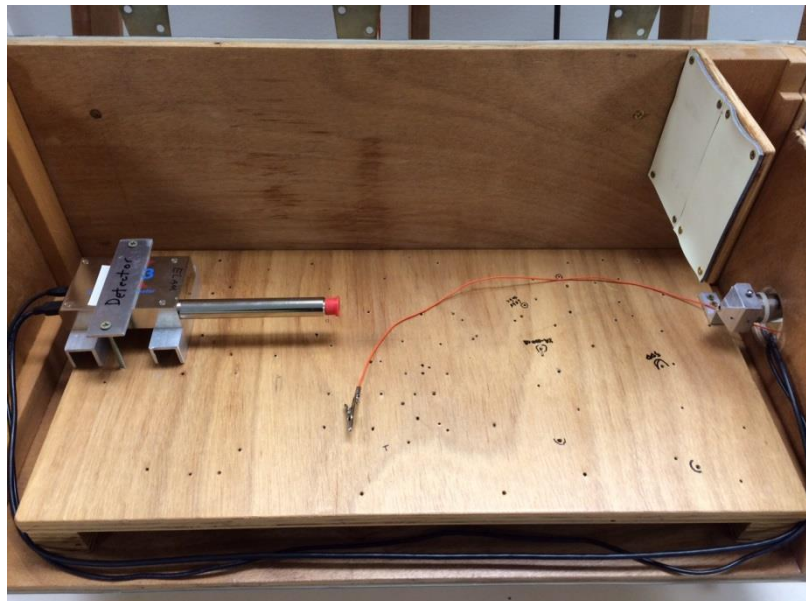


- Number of electrons contributing to the Compton scattering response calculated numerically
- Energy of each electron is calculated
- Compton response and source peak response are summed as the total response function

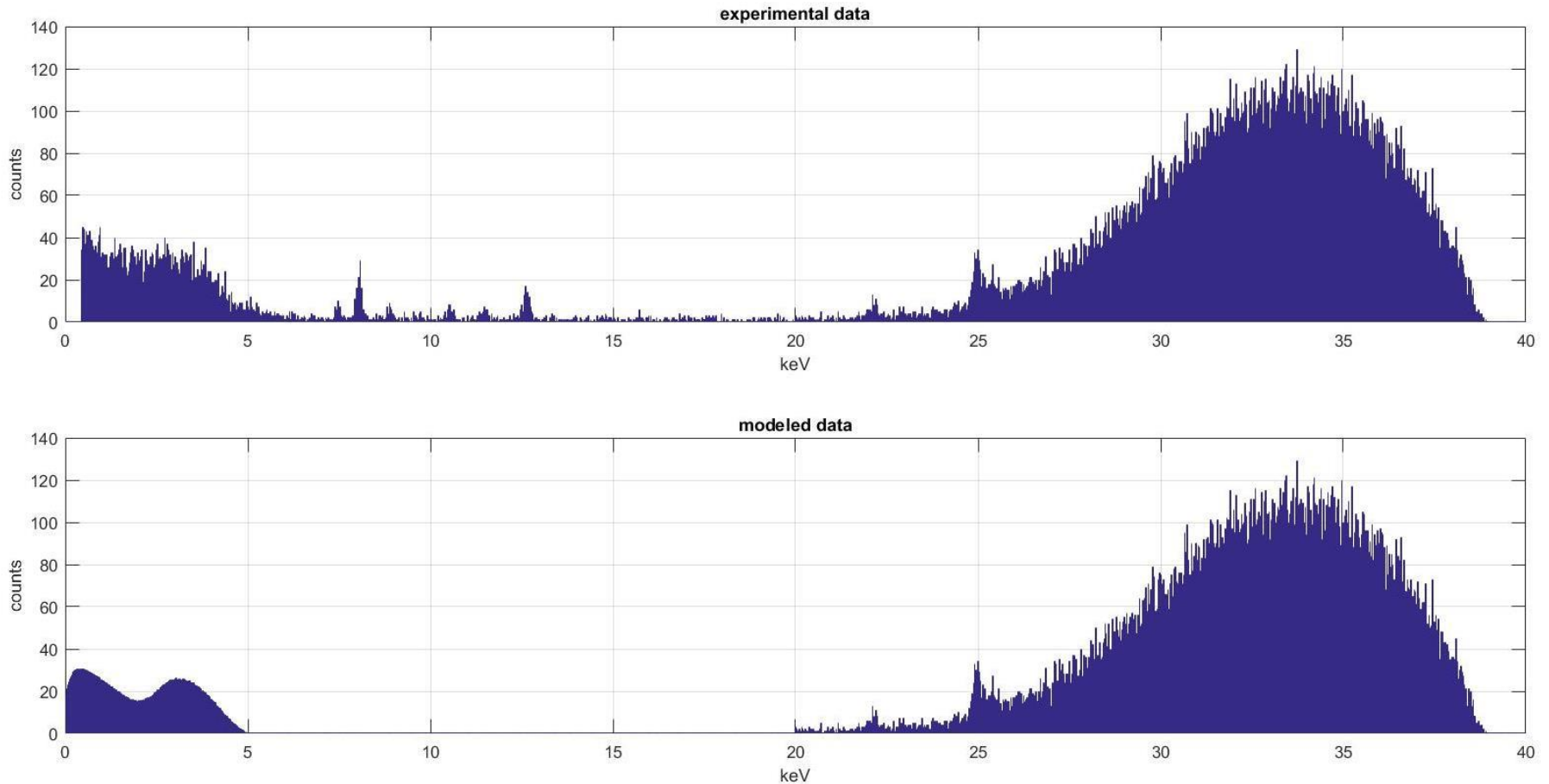
$$N_{scattered\_electrons} = \sum_{E=0}^{E=E_{max}} \sum_{D=0}^{D=detector\_thickness} N_{0,E} e^{(-\sigma_{incident\_photon}\rho D)} \sum_{\phi=0}^{\phi=\pi} \left( 1 - e^{\left(-\frac{d\sigma_c}{d\phi} d\phi_{atomic} dD\right)} \right) \left( e^{(-\sigma_{escape\_photon}\rho EL)} \right) d\phi dD dE$$

$$E_{electron} = E_0 \frac{\alpha(1 - \cos \phi)}{1 + \alpha(1 - \cos \phi)}$$

# Experimental Data



# Comparison of model and experiment



Looks pretty good!

# What's next?



- Design and run a better experiment: higher counts, cleaner data
- Statistical comparison of model and experimental data to quantify the accuracy of the model
- Send this thing to Mars and check out some rocks

Thank you  
&  
any questions???

